## Lecture 4:

# Synchrotron Radiation 

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## Lienard-Wiechert Formula

$$
\begin{aligned}
& \text { Space Charge } \\
& \downarrow \\
& \vec{E}=e\left[\frac{\vec{n}-\vec{\beta}}{\gamma^{2}(1-\vec{n} \cdot \vec{\beta})^{3} R^{2}}\right]_{r e t}+\left(\frac{e}{c}\right)\left[\frac{\vec{n} \times(\vec{n}-\vec{\beta}) \times \dot{\vec{\beta}}}{(1-\vec{n} \cdot \vec{\beta})^{3} R}\right]_{r e t}, \\
& \vec{B}=\vec{n} \times \vec{E}
\end{aligned}
$$

* Space charge is suppressed by $1 / \gamma^{2}$
* Identify radiated field with synchrotron radiation
* Subject to retarded condition:

$$
t^{\prime}=t-\frac{R\left(t^{\prime}\right)}{c}
$$

## Spectral Distribution

In the far-field approximation, the intensity distribution is given by,

$$
\frac{d^{2} I}{d \omega d \Omega}=\frac{r_{e} m c \omega^{2}}{4 \pi^{2}}\left|\int_{-\infty}^{\infty} \hat{n} \times\left[\hat{n} \times \vec{\beta}\left(t^{\prime}\right)\right] \exp \left[i \omega\left(t^{\prime}-\hat{n} \cdot \vec{r}\left(t^{\prime}\right) / c\right)\right] d t^{\prime}\right|^{2}
$$



## Computing Radiation Spectrum



Radiation direction:

$$
\hat{n}=(0, \sin \theta, \cos \theta)
$$

Electron position:

$$
\left.\vec{r}(t)=\left(-\rho\left(1-\cos \left(\frac{\nu t}{\rho}\right)\right), 0, \rho \sin \left(\frac{\nu t}{\rho}\right)\right)\right)
$$

Its velocity:

$$
\vec{\beta}(t)=\left(-\beta \sin \left(\frac{\nu t}{\rho}\right), 0, \beta \cos \left(\frac{v t}{\rho}\right)\right)
$$

Phase approximation:

$$
\omega\left(t-\frac{\hat{n} \cdot \vec{r}(t)}{c}\right)=\omega\left[t-\frac{\rho}{c} \sin \left(\frac{\nu t}{\rho}\right) \cos \theta\right] \approx \frac{\omega}{2}\left[\left(\frac{1}{\gamma^{2}}+\theta^{2}\right) t+\frac{c^{2}}{3 \rho^{2}} t^{3}\right]
$$

Vector integrand:

$$
\hat{n} \times(\hat{n} \times \vec{\beta})=\beta\left[\sin \left(\frac{v t}{\rho}\right) \hat{\varepsilon}_{\sigma}+\cos \left(\frac{\nu t}{\rho}\right) \sin \theta \hat{\varepsilon}_{\pi}\right] \approx \frac{c t}{\rho} \hat{\varepsilon}_{\sigma}+\theta \hat{\varepsilon}_{\pi}
$$

## Radiation Spectrum by Bending Magnet

Intensity distribution is given by, $\begin{array}{ll}\sigma \text { mode } & \pi \text { mode } \\ \downarrow & \downarrow\end{array}$

$$
\frac{d^{2} I}{d \omega d \Omega}=\frac{3 r_{e} m c}{4 \pi^{2}} \gamma^{2}\left(\frac{\omega}{\omega_{c}}\right)^{2}\left(1+\gamma^{2} \theta^{2}\right)^{2}\left[K_{2 / 3}^{2}(\xi)+\frac{\gamma^{2} \theta^{2}}{1+\gamma^{2} \theta^{2}} K_{1 / 3}^{2}(\xi)\right]
$$

where $\mathrm{K}_{1 / 3}$ and $\mathrm{K}_{2 / 3}$ are modified Bessel functions and their argument

$$
\xi=\frac{1}{2} \frac{\omega}{\omega_{c}}\left(1+\gamma^{2} \theta^{2}\right)^{3 / 2}
$$

Angle integrated intensity distribution is

$$
\frac{d I}{d \omega}=\sqrt{3} r_{e} m c \gamma \frac{\omega}{\omega_{c}} \int_{\omega / \omega_{c}}^{\infty} K_{5 / 3}(x) d x
$$

where the critical frequency is

$$
\omega_{c}=\frac{3}{2} \gamma^{3}\left(\frac{c}{\rho}\right)
$$

# Intensity Distribution 

o mode

$\pi$ mode



## Radiation Spectrum



## Beam Dynamics in Undulator

Electron velocity:
Its position:

$$
\begin{aligned}
& \frac{d x}{d t}=-\beta c \frac{K}{\gamma} \sin \left(k_{p} z\right) \\
& \frac{d z}{d t}=\beta c\left[1-\frac{K^{2}}{2 \gamma^{2}} \sin ^{2}\left(k_{p} z\right)\right]
\end{aligned}
$$

$$
\begin{aligned}
& x(t)=\frac{K}{\gamma k_{p}} \cos \left(k_{p} \bar{\beta} c t\right) \\
& z(t)=\bar{\beta} c t+\frac{K^{2}}{8 \gamma^{2} k_{p}} \sin \left(2 k_{p} \bar{\beta} c t\right)
\end{aligned}
$$

where the undulator parameter K and averaged velocity:

$$
\begin{aligned}
& K=\frac{e B_{0} \lambda_{p}}{2 \pi m c^{2}} \\
& \bar{\beta}=\beta\left(1-\frac{K^{2}}{4 \gamma^{2}}\right)
\end{aligned}
$$

## Computing Spectrum of Undulator Radiation

Radiation direction:


$$
\hat{n}=(\cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta)
$$

Electron position:

$$
\vec{r}(t)=\left(\frac{K}{k_{p} \gamma} \cos \left(\omega_{p} t\right), 0, \bar{\beta} c t+\frac{K^{2}}{8 \gamma^{2} k_{p}} \sin \left(2 \omega_{p} t\right)\right)
$$

Its velocity:
beam
Phase approximation:

$$
\vec{\beta}(t)=\left(-\frac{K}{\gamma} \sin \left(\omega_{p} t\right), 0, \bar{\beta}\left[1+\frac{K^{2}}{4 \gamma^{2}} \cos \left(2 \omega_{p} t\right)\right]\right)
$$

$$
\omega\left(t-\frac{\hat{n} \cdot \vec{r}(t)}{c}\right) \approx \frac{\omega}{\omega_{1}}\left[\omega_{p} t-\frac{K \bar{\beta} \theta}{\gamma} \frac{\omega_{1}}{\omega_{p}} \cos \phi \cos \left(\omega_{p} t\right)-\frac{K^{2} \bar{\beta}}{8 \gamma^{2}} \frac{\omega_{1}}{\omega_{p}} \sin \left(2 \omega_{p} t\right)\right]
$$

Vector integrand:

$$
\hat{n} \times(\hat{n} \times \vec{\beta}) \approx \bar{\beta}\left\{\left[\theta \cos \phi+\frac{K}{\gamma} \sin \left(\omega_{p} t\right)\right] \hat{x}+\theta \sin \phi \hat{y}\right\}
$$

## Number of Photons within $\Delta \omega / \omega$

Total emitted photons after an electron passing through undulator is given by,

$$
\frac{d N_{p h}(\omega)}{d \Omega}=\alpha \gamma^{2} \bar{\beta}^{2} N_{p}^{2} \frac{\Delta \omega}{\omega} \sum_{k=1}^{\infty} k^{2}\left[\frac{\sin \left(\pi N_{p} \Delta \omega_{k} / \omega_{1}\right)}{\pi N_{p} \Delta \omega_{k} / \omega_{1}}\right]^{2}\left[I_{\sigma, k}+I_{\pi, k}\right]
$$

where,

$$
\begin{aligned}
& I_{\sigma, k}=\frac{\left(2 \gamma \theta \Sigma_{1} \cos \phi-K \Sigma_{2}\right)^{2}}{\left(1+\frac{K^{2}}{2}+\gamma^{2} \theta^{2}\right)^{2}} \\
& I_{\pi, k}=\frac{\left(2 \gamma \theta \Sigma_{1} \sin \phi\right)^{2}}{\left(1+\frac{K^{2}}{2}+\gamma^{2} \theta^{2}\right)^{2}}
\end{aligned}
$$

and

$$
\begin{aligned}
& \Sigma_{1}=\sum_{m=-\infty}^{\infty} J_{-m}(\mu) J_{k-2 m}(v) \\
& \Sigma_{2}=\sum_{m=-\infty}^{\infty} J_{-m}(\mu)\left[J_{k-2 m-1}(v)+J_{k-2 m+1}(v)\right] \\
& \text { and } J_{n} \text { are Bessel functions. }
\end{aligned}
$$

## Interference Spectrum



$$
\frac{\Delta \omega_{k}}{\omega_{k}}= \pm \frac{1}{k N_{p}} \quad \text { first zeros near the origin define the width of the peak. }
$$

# Radiation Distribution of $\sigma$ and $\pi$ Modes $(K=1)$ 








## Forward Radiation

Total emitted photons after an electron passing through undulator is given by,

$$
\frac{d N_{p h}(\omega)}{d \Omega}=\alpha \gamma^{2} N_{p}^{2} \frac{\Delta \omega}{\omega} \sum_{k=1}^{\infty} A_{k}(K)\left[\frac{\sin \left(\pi N_{p} \Delta \omega_{k} / \omega_{1}\right)}{\pi N_{p} \Delta \omega_{k} / \omega_{1}}\right]^{2}
$$

where

$$
\begin{aligned}
& A_{k}(K)=\frac{k^{2} K^{2}}{\left(1+\frac{K^{2}}{2}\right)^{2}} J J^{2} \\
& J J=\left[J_{(k+1) 2}\left(\frac{k K^{2}}{4+2 K^{2}}\right)-J_{(k-1) 2}\left(\frac{k K^{2}}{4+2 K^{2}}\right)\right]
\end{aligned}
$$

Only odd $\sigma$ modes contribute


Undulator parameter K should be between 1 to 4

## Photon Flux

Flux at $\mathrm{k}^{\text {th }}$ harmonics:

$$
\left.\frac{d N_{p h}\left(\omega_{k}\right)}{d t}\right|_{\theta=0}=\frac{\pi}{2} \alpha N_{p} \frac{I}{e} \frac{\Delta \omega}{\omega_{k}} Q_{k}(K)
$$

where

$$
Q_{k}(K)=\frac{1+\frac{K^{2}}{2}}{k} A_{k}(K)
$$

The rms opening angle:


$$
\sigma_{r^{\prime}} \approx \frac{1}{2 \gamma} \sqrt{\frac{1+\frac{K^{2}}{2}}{k N_{p}}}=\sqrt{\frac{\lambda_{k}}{2 L}}
$$

Undulator parameter: K has to be large enough

The forward cone: $\quad d \Omega=2 \pi \sigma_{r^{\prime}}^{2}$

## Photon Flux of PEP-X


$\mathrm{n}^{\text {th }}$ harmonic wavelength:

$$
\lambda_{n}=\frac{\lambda_{u}}{2 n \gamma^{2}}\left(1+\frac{K^{2}}{2}\right) \quad F_{n}=\frac{\pi}{2} \alpha N_{u} Q_{n}\left(\frac{n K^{2}}{4+2 K^{2}}\right) \frac{\Delta \omega}{\omega} \frac{I}{e}
$$

## Gaussian Mode

The fundamental Gaussian mode can be written as

$$
E(x, y, z)=E_{0} \frac{w_{0}}{w(z)} \exp \left[-\frac{r^{2}}{w(z)}\right] \exp \left[-i\left(k z+k \frac{r^{2}}{2 R(z)}-\phi(z)\right)\right]
$$

where
spot size: $\quad w(z)=w_{0} \sqrt{1+\left(z / z_{R}\right)^{2}}$
radius of curvature: $R(z)=z\left[1+\left(z / z_{R}\right)^{2}\right]$
Guoy phase:

$$
\phi(z)=\tan ^{-1}\left(z / z_{R}\right)
$$

Rayleigh length: $\quad z_{R}=\frac{\pi w_{0}^{2}}{\lambda}$
It is a solution of the paraxial wave equation:

$$
\left(\nabla_{\perp}^{2}-2 i k \frac{\partial}{\partial z}\right) \psi(x, y, z)=0
$$

## Visualization of a Gaussian Mode



## Brightness of Gaussian Mode

For a Gaussian mode, its brightness distribution function is given by,

$$
\begin{aligned}
B(\vec{r}, \vec{\varphi} ; 0) & =B_{0} \exp \left[-\frac{\vec{r}^{2}}{2 \sigma_{r}^{2}}-\frac{\vec{\varphi}^{2}}{2 \sigma_{r^{\prime}}^{2}}\right] \\
\sigma_{r} & =w_{0} / 2 \\
\sigma_{r^{\prime}} & =\sigma_{r} / z_{R}
\end{aligned}
$$

Then, we have

$$
\begin{gathered}
\sigma_{r} \sigma_{r^{\prime}}=\lambda / 4 \pi \quad \text { emittance } \\
\sigma_{r} / \sigma_{r^{\prime}}=z_{R} \quad \text { beta function } \\
B_{0}=\frac{F}{\left(2 \pi \sigma_{r} \sigma_{r^{\prime}}\right)^{2}}=\frac{F}{(\lambda / 2)^{2}} \longleftarrow \text { coherence volume }
\end{gathered}
$$

## Single Electron Brightness

Using the Gaussian mode as an approximation for the undulator source, we choice $z_{R}=L / 2 \pi$, so that,

$$
\begin{aligned}
& \sigma_{r^{\prime}}=\sqrt{\frac{\lambda_{k}}{2 L}} \\
& \sigma_{r}=\frac{\sqrt{2 \lambda_{k} L}}{4 \pi}
\end{aligned}
$$

Its brightness function is given by,

$$
B(\vec{r}, \vec{\varphi} ; 0)=B_{0} \exp \left[-\frac{\vec{r}^{2}}{2 \sigma_{r}^{2}}-\frac{\vec{\varphi}^{2}}{2 \sigma_{r^{\prime}}^{2}}\right]
$$

and the photon flux is

$$
F=\frac{\pi}{2} \alpha N_{p} \frac{I}{e} \frac{\Delta \omega}{\omega_{k}} Q_{k}(K)
$$

## Spectral Brightness of Electron Beam

Brightness of electron beam radiating at $\mathrm{n}^{\text {th }}$ (odd) harmonics in a undulator is given by

$$
B_{k}=F_{k} /\left(4 \pi^{2} \Sigma_{x} \Sigma_{x}^{\prime} \Sigma_{y} \Sigma_{y}^{\prime}\right)
$$

If the electron beam phase the brightness becomes optimized

$$
B_{k}=\frac{F_{k}}{4 \pi^{2}\left(\varepsilon_{x}+\lambda_{k} / 4 \pi\right)\left(\varepsilon_{y}+\lambda_{k} / 4 \pi\right)}
$$

space is matched to those of photon's,

Finally, even for zero emittances, there is an ultimate limit for the brightness

Spectral brightness of PEP-X


$$
B_{k}=\frac{4 F_{k}}{\lambda_{k}^{2}}
$$

A diffraction limited ring at 1 angstrom or 8 pm-rad emittance

## Coherent X-Ray Diffraction Imaging with nanofocused Illumination C.G. Schroer et al. PRL 101, 090801 (2008)

- Phone energy: 15.25 keV
- Coherent flux: $10^{8} \mathrm{ph} / \mathrm{s}$
- Exposure time: $60 x 10$ s
- Resolution: 5 nm
- $\Delta E / E: 1.4 \times 10^{-4}$

The total number of photons $D_{c}$ in the coherence volume available at a given source, however, is bounded from above by

$$
D_{c}=F_{c} T=\operatorname{Br} \lambda^{2} \frac{\Delta E}{E} T
$$

where $F_{c}$ is the coherent flux, Br is the brilliance of the x-ray source, $\lambda$ is the wavelength of the x rays, $\Delta E / E$ the degree of monochromaticity, and $T$ the exposure time. For

## Improvement of resolution scaled as $D_{c}{ }^{1 / 4}$.



FIG. 1 (color online). (a) Schematic sketch of the coherent diffraction imaging setup with nanofocused illumination. (b) Scanning electron micrograph of gold particles (diameter $\approx 100 \mathrm{~nm}$ ) deposited on a $\mathrm{Si}_{3} \mathrm{~N}_{4}$ membrane. (c) Diffraction pattern (logarithmic scale) recorded of the single gold particle pointed to by the arrow in (b) and illuminated by a hard x-ray beam with lateral dimensions of about $100 \times 100 \mathrm{~nm}^{2}$. The maximal momentum transfer, both in horizontal and vertical direction, is $q=1.65 \mathrm{~nm}^{-1}$.

## THE DEGREE OF TEMPORAL COHERENCE IS DETERMINED BY THE LENGTH OF THE WAVE TRAIN (MONOCHROMATICITY)



- The main point is to make sure that the coherence length is long compared to all path differences between interfering rays in the experiment
- If this is done then the illumination is called quasimonochromatic and temporal coherence effects are removed from consideration


## THE DEGREE OF SPATIAL COHERENCE IS DETERMINED BY THE DEGREE OF COLLIMATION



## YOUNG'S SLITS EXPERIMENT IN COHERENT ILLUMINATION



The fringe visibility $\mathbf{V}$ is given by

$$
\mathbf{V}=\frac{I_{\max }-I_{\min }}{I_{\max }+I_{\min }}=\frac{4 \sqrt{I_{1} I_{2}}}{2\left(I_{1}+I_{2}\right)}=1 \text { when } I_{1}=I_{2}
$$

## INTRODUCE A TILTED WAVE TO REPRESENT IMPERFECT COLLIMATION



## FRINGE CONTRAST WHEN THE ILLUMINATING BEAM HAS ANGULAR SPREAD

- The graphs show the loss of fringe contrast when the fringe patterns with all $\delta$ values in the given $\delta$ range were averaged from $-\delta / 2$ to $+\delta / 2$
- $\delta$ range equals zero is the coherent case
- Note that there is no change in the phase of the fringes because the angular spread was symmetrical
- The zero and maximum of the intensity for each plotted fringe pattern are the axes immediately above and below the plot
- When $\delta$ equals one wavelength for example the total beam angular spread is $1 \lambda / \mathrm{P}_{1} \mathrm{P}_{2}$


## THE UNDULATOR ONE-ELECTRON PATTERN

- The on-axis monochromatic one-electron pattern emitted by an undulator is a spatially-coherent beam - also known as a diffraction-limited beam or a wave mode
- We will model it as a Gaussian laser mode with RMS intensity width and angular width equal to $\sigma_{r}$ and $\sigma_{r^{\prime}}$ - so that the width-angle product or emittance is given by

$$
\sigma_{r} \sigma_{r^{\prime}}=\frac{\lambda}{4 \pi}
$$

- We will rearrange this using the fact that a rectangle of width $\sqrt{2 \pi} \sigma$ and height 1 has equal area to a Gaussian of RMS width $\sigma$ and height 1 - thus we get

$$
\begin{aligned}
\left(\sqrt{2 \pi} \sigma_{r}\right)\left(\sqrt{2 \pi} \sigma_{r^{\prime}}\right)=\frac{\lambda}{2} & \text { Worth } \\
& \Delta_{c} \Delta_{c}^{\prime}=\frac{\lambda}{2},
\end{aligned}
$$



- Where $\Delta_{c}=\sqrt{2 \pi} \sigma_{r}$ and $\Delta_{c}^{\prime}=\sqrt{2 \pi} \sigma_{r^{\prime}}$, this is the relation you use to choose beam-line slit widths to get a coherent beam
- This is now the same as our earlier representation of a spatially coherent beam

$$
a A=\frac{\lambda}{2}
$$

## References

1) J.D. Jackson, Classical Electrodynamics, Third Edition, John Wiley \& Son, Inc. 1999
2) H. Wiedemann, Synchrotron Radiation, Springer-Verlag Berlin Heidelberg 2003
3) Kwang-Je Kim, "Characteristics of Synchrotron Radiation," AIP Proc. No. 184 (AIP, New York, 1989), pp. 565-632
4) Malcolm Howells, ESRF lecture series of coherent X-ray and their applications
