



Stellar Interferometer - Theory and Practice

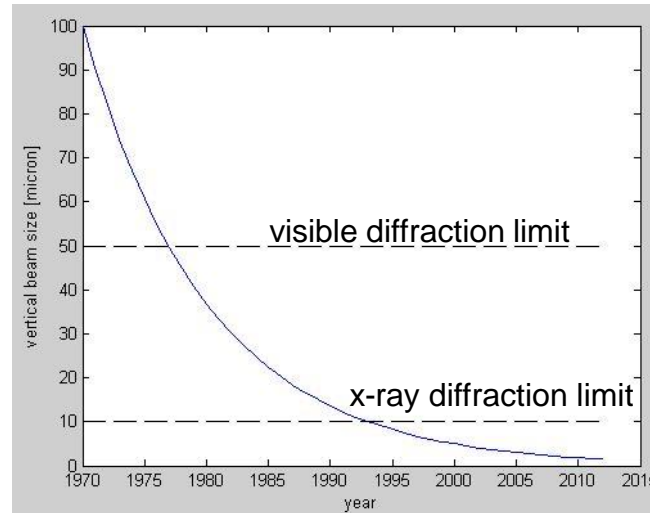
US Particle Accelerator School
January 18-22, 2010

- Motivation
- Two-slit Interference – Young's experiment
- Diffraction from a single slit - review
- Extended Source – Partial Coherence
- The Mutual Coherence function
- Van-Cittert/Ziernike theorem
- Stellar Interferometers for SR applications



Motivation for Interferometry

- Electron beam size can be very small



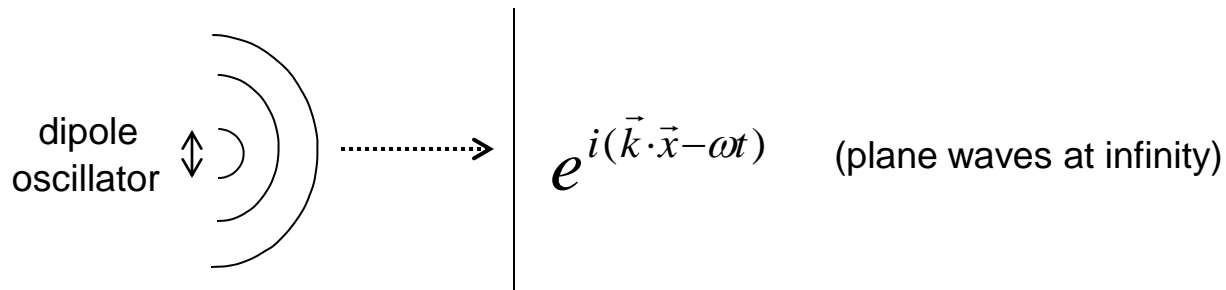
$$\sqrt{\varepsilon\beta} = \sqrt{(1)(1 \times 10^{-9} / 1000)} = 1 \mu m!$$

- Need to measure beam size for optics verification, machine monitoring and operation
- Conventional imaging diffraction limited
 - $\sigma_{\text{res}} \sim 50 \text{ um}$ visible
 - $\sigma_{\text{res}} \sim 10 \text{ um}$ x-ray pinhole
- What else can be used?



Motivation for Interferometry (cont'd)

- Take advantage of light *coherence* properties



- For a *distributed* source at finite distance the light is only *partially coherent*

$$\vec{k} = k_x \hat{x} + k_z \hat{z}$$

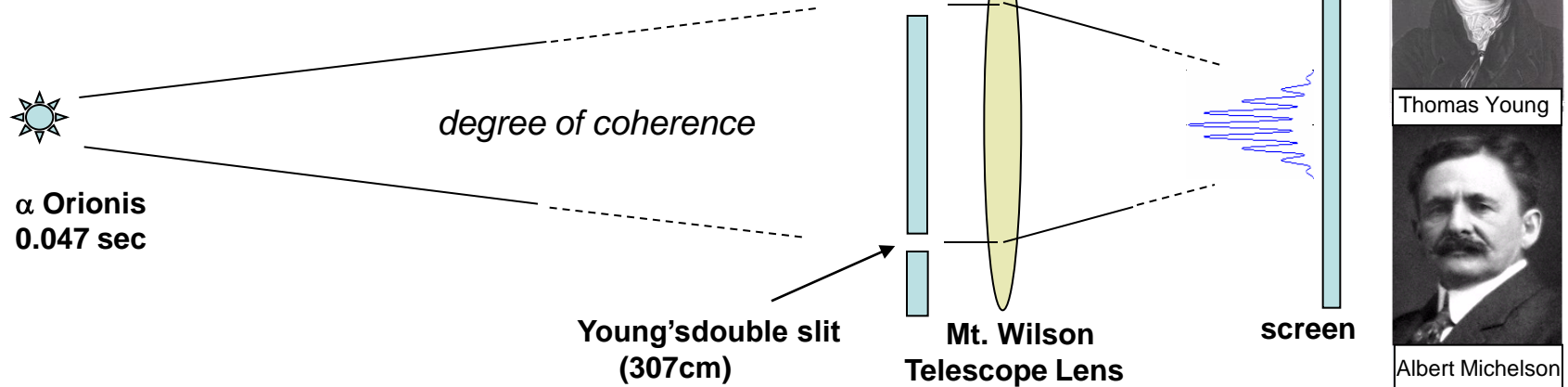
The diagram shows a bundle of light rays that are slightly diverging. To the right of the rays, the wave vector \vec{k} is expressed as the sum of its components: $k_x \hat{x} + k_z \hat{z}$. The k_x term is enclosed in a dashed oval.

- Interferometry enters world of wavefront physics and statistical optics

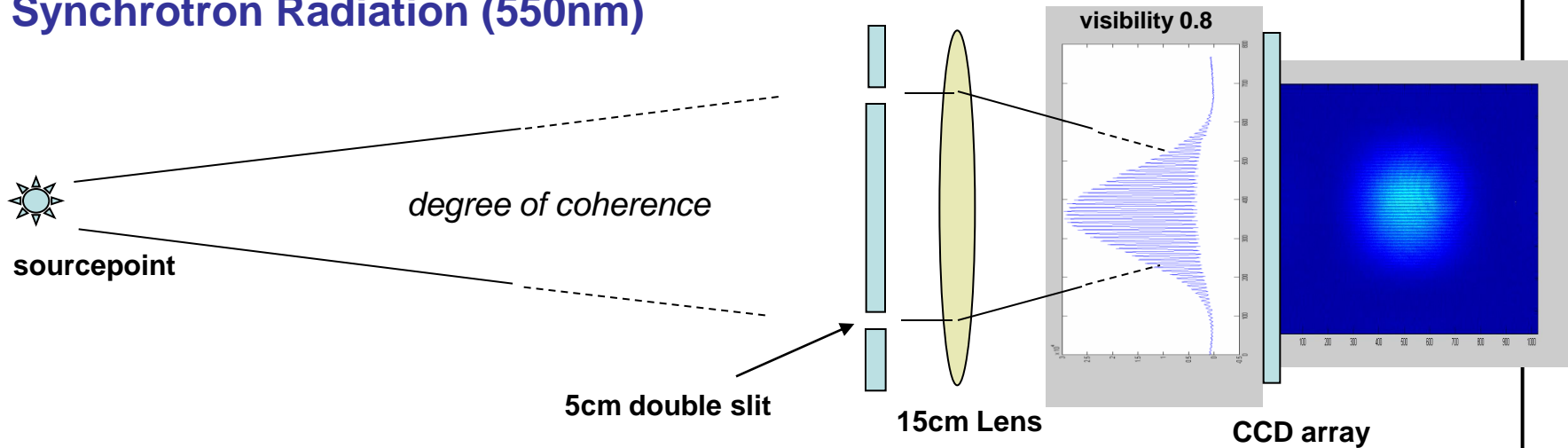


Interferometric Beam Size Measurement

Michelson (circa 1890)



Synchrotron Radiation (550nm)





Two-Slit Interference: Young's Experiment

Diagram illustrating the setup for Young's double-slit experiment. Two slits are separated by a distance d . Light rays are shown emerging from the slits at an angle θ . The path difference between the rays is given by the equation:

$$d \sin \theta = m \frac{\lambda}{2}$$

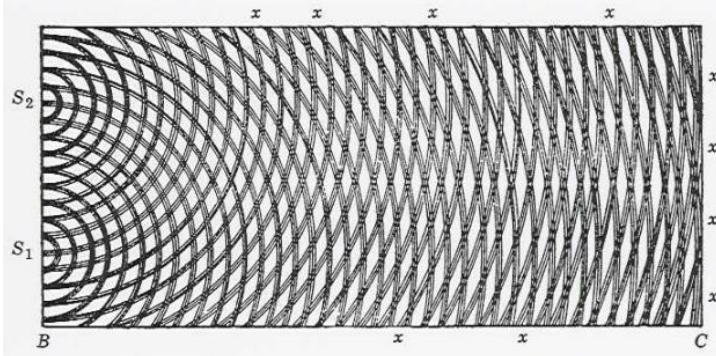
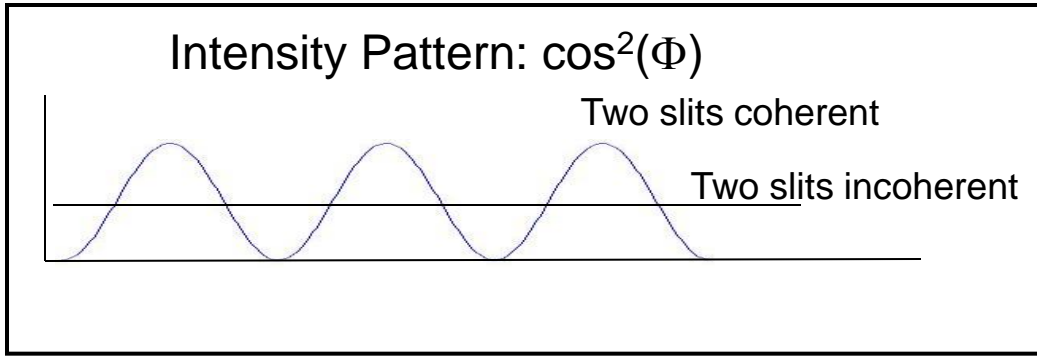
The resulting field pattern on the screen is described by the equation:

$$E = E_o e^{i(\Phi)} + E_o e^{i(-\Phi)} = 2E_o \cos(\Phi)$$

The interference pattern is represented by the integral:

$$\int_{-\infty}^{\infty} (\delta(\Phi) + \delta(-\Phi)) e^{ikx} dx$$

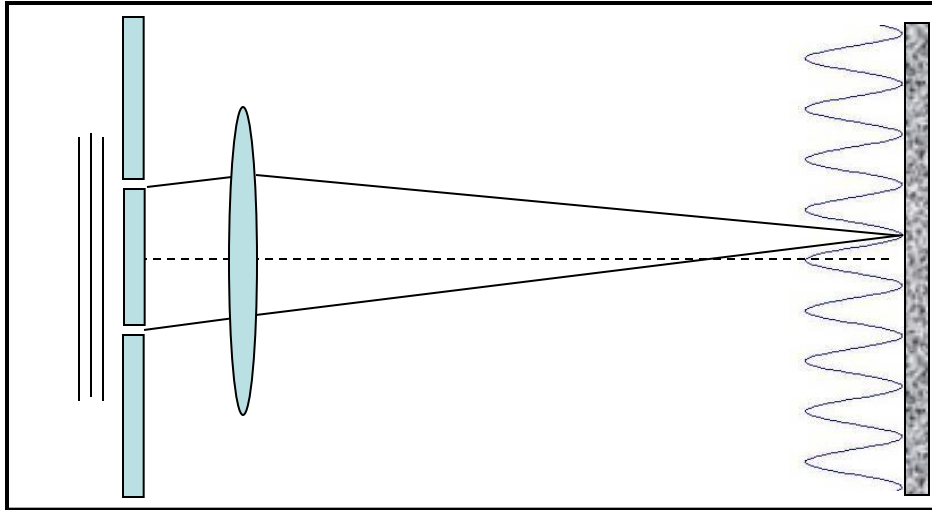
monochromatic light!





Two-Slit Interference (cont'd)

Use a lens to concentrate image on screen



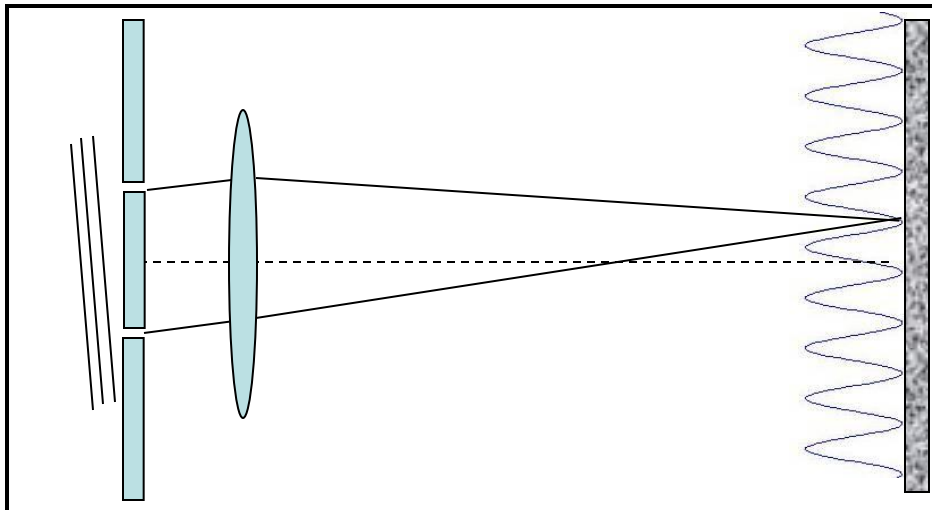
Intensity Pattern

Note Contrast/Visibility

$$V = \frac{I_{Max} - I_{Min}}{I_{Max} + I_{Min}} = 1.0$$

fully coherent!

Change phase/incidence angle, shift pattern phase

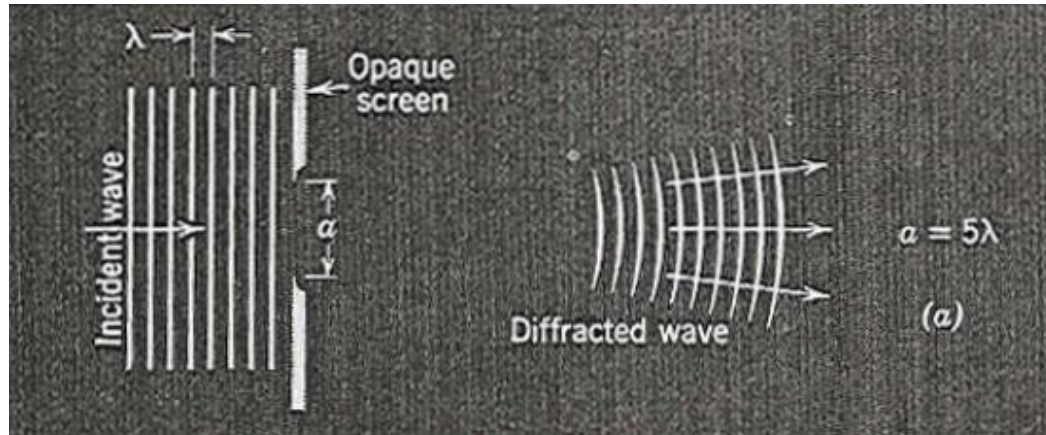


$V=1.0$ (no change)

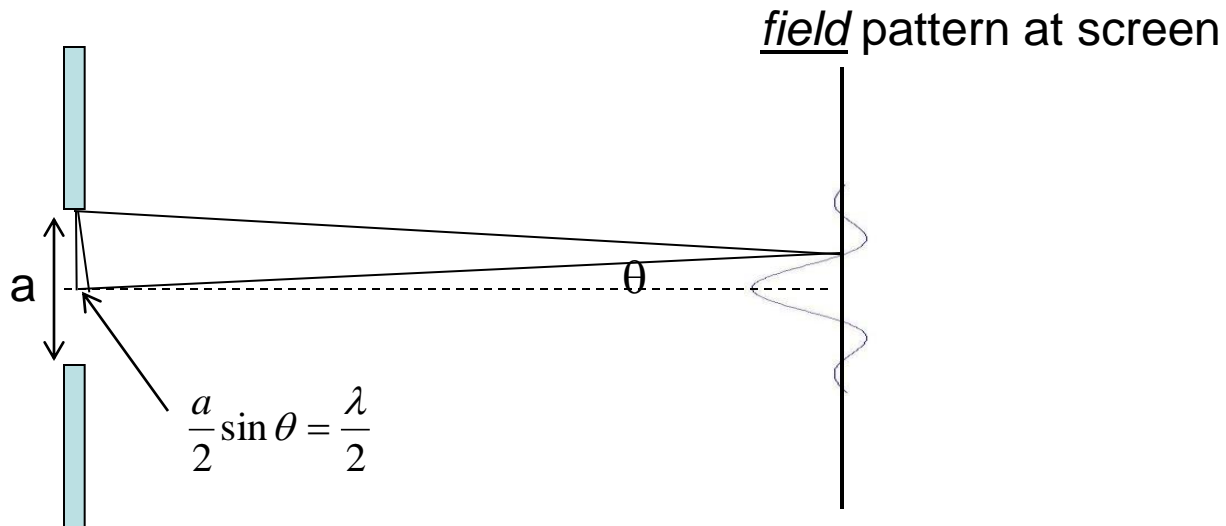


Single-Slit Diffraction - Review

Plane wave incident on aperture - diffraction

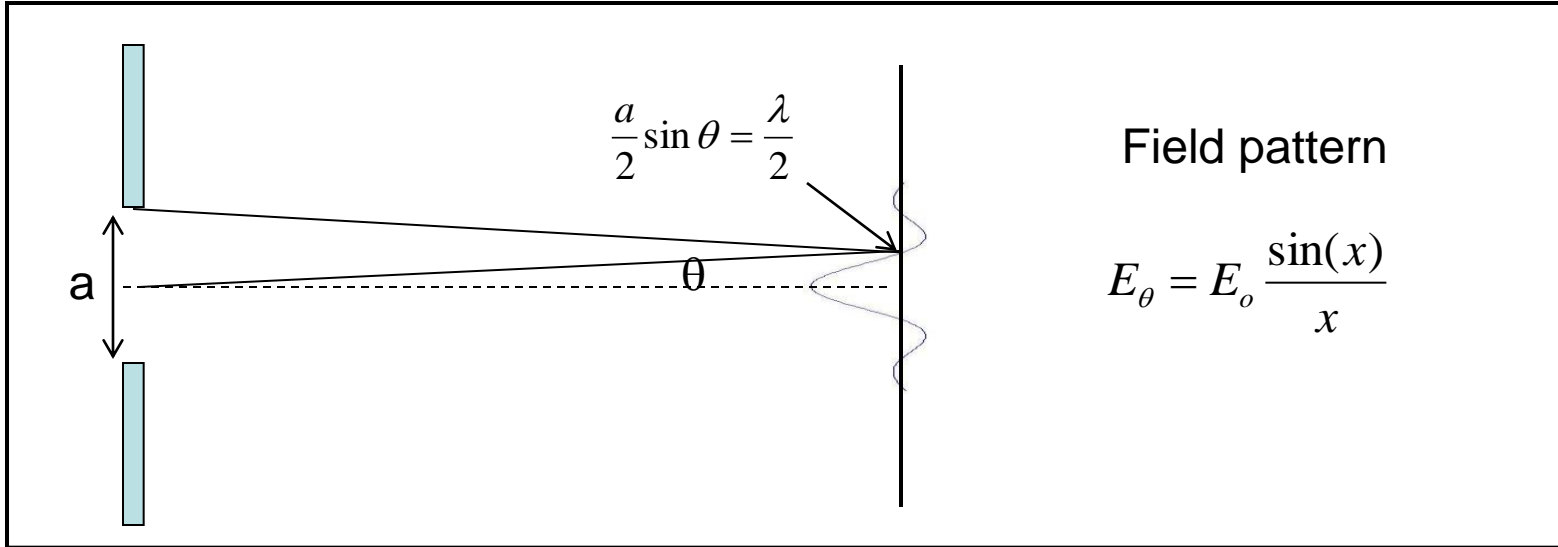


Condition for first diffraction minima



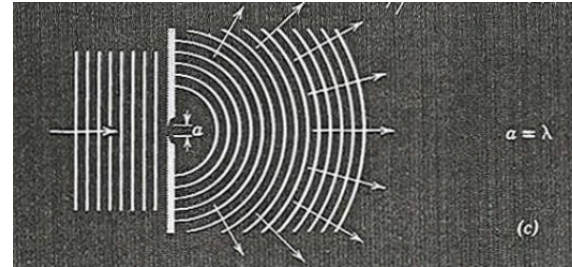
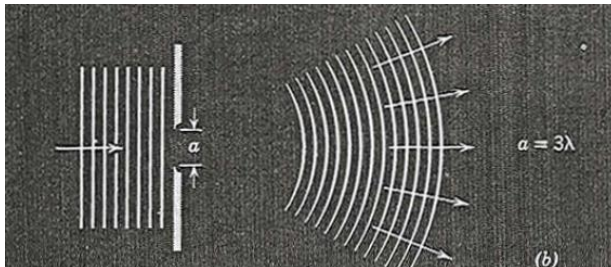


Single-Slit: Electric Field Pattern



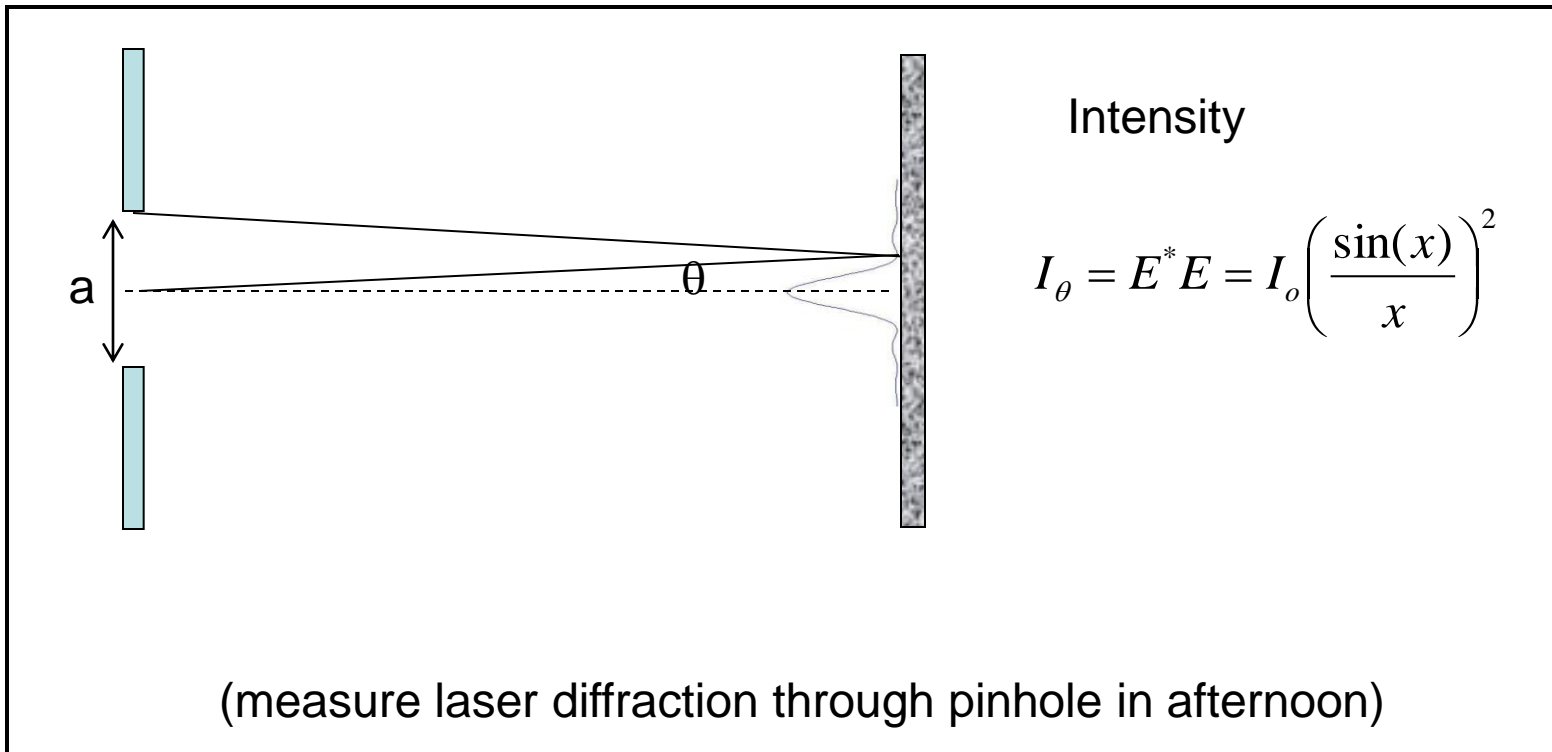
Field pattern is the Fourier Transform of the Aperture

$$\int_{-x_o}^{x_o} e^{ikx} dx = \frac{e^{ikx_o} - e^{-ikx_o}}{ik} = \frac{\sin(kx_o)}{k}$$





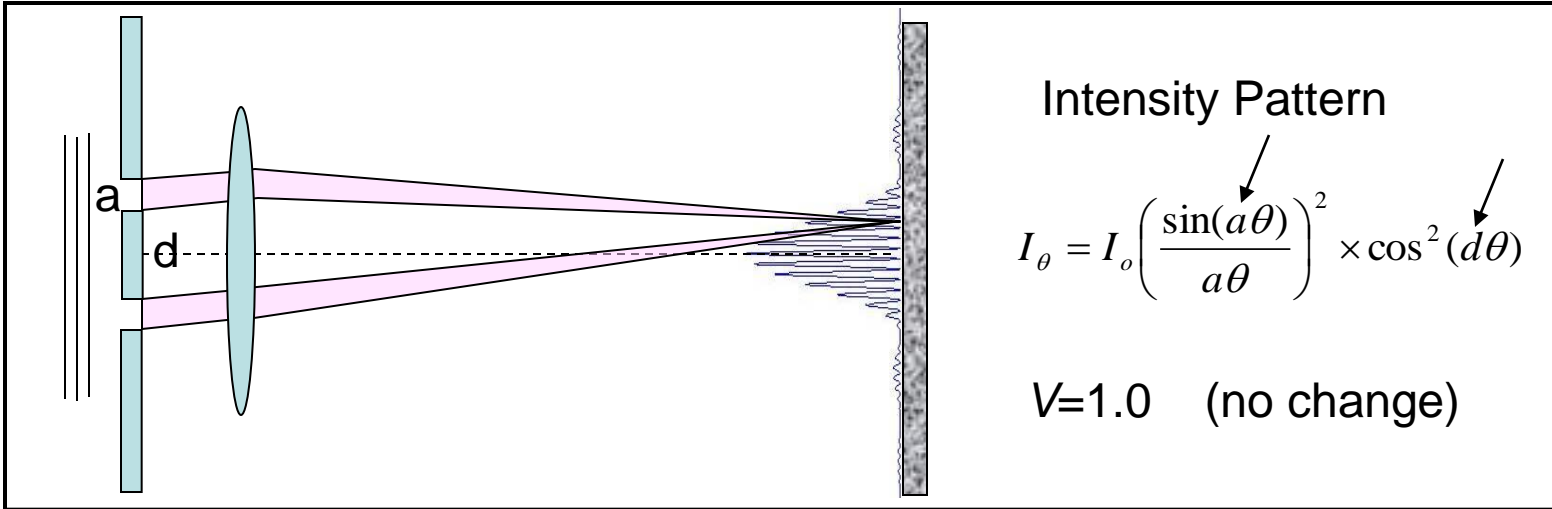
Single-Slit: Intensity Pattern





Two-Slit Interference (cont'd)

Consider 2-slit interference with finite slit size:



Important mathematical point: $1 + \cos(\theta) = 2 \cos^2\left(\frac{\theta}{2}\right)$

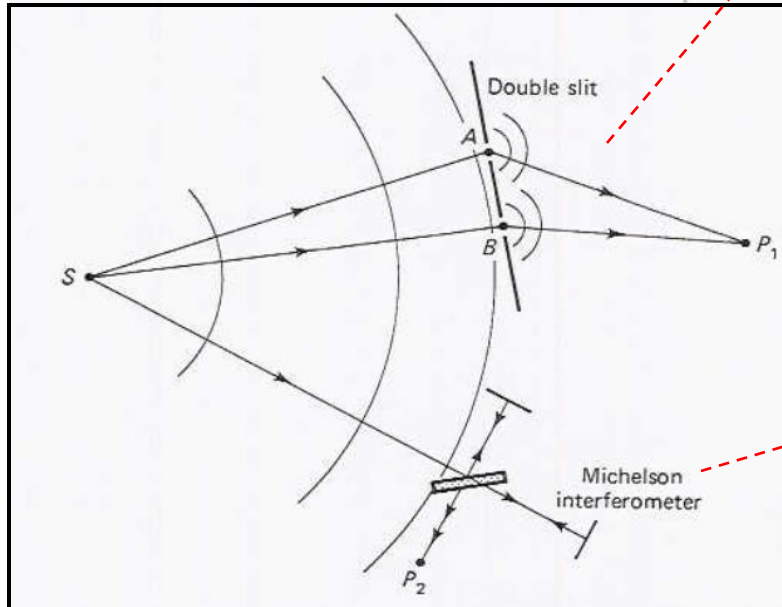
Intensity pattern can be written: $I_{\theta} = I_o \underbrace{\left(\frac{\sin(a\theta)}{a\theta} \right)^2}_{\text{Single-Slit}} \times \underbrace{(1 + \cos(d\theta))}_{\text{Two-Slit}}$

This is *approximately* the form we will work with (but visibility 1.0)



From Interference to Interferometry

Interferometry is used to measure *coherence* properties of light



Spatial Coherence

coherence length is inversely proportional to size

*a star emits plane-waves randomly in direction
-emission is not coherent at close distance*

*a point source emits spherical-waves in all directions
-emission is coherent at long distances*

Temporal Coherence

coherence time is inversely proportional to line-width

*thermal source emits wavepackets randomly in time
-emission is not coherent*

*laser source emits continuous wavetrain in time
-emission is coherent*



Degree of Coherence

Light can be totally coherent, partially coherent or incoherent

The degree of coherence is found from a correlation function

Temporal Degree of Coherence

Consider two colinear light waves $E_1(t)$ and $E_2(t)$

the correlation function is $\Gamma_{12}(\tau) = \langle E_1(t)E_2^*(t + \tau) \rangle$

Γ_{12} is the degree of self-coherence

if $\tau < \tau_0$ (correlation time), then waves are coherent
and light interferes

if $\tau > \tau_0$ then coherence is lost
and light waves do not interfere

Spatial Degree of Coherence

Consider two waves $E_1(\vec{k})$ and $E_2(\vec{k})$ from two sources

the correlation function is now $\Gamma_{12}(r) = \langle E_1(P_1)E_2^*(P_2) \rangle$

P_1 and P_2 are two points in space and $r = P_1 - P_2$

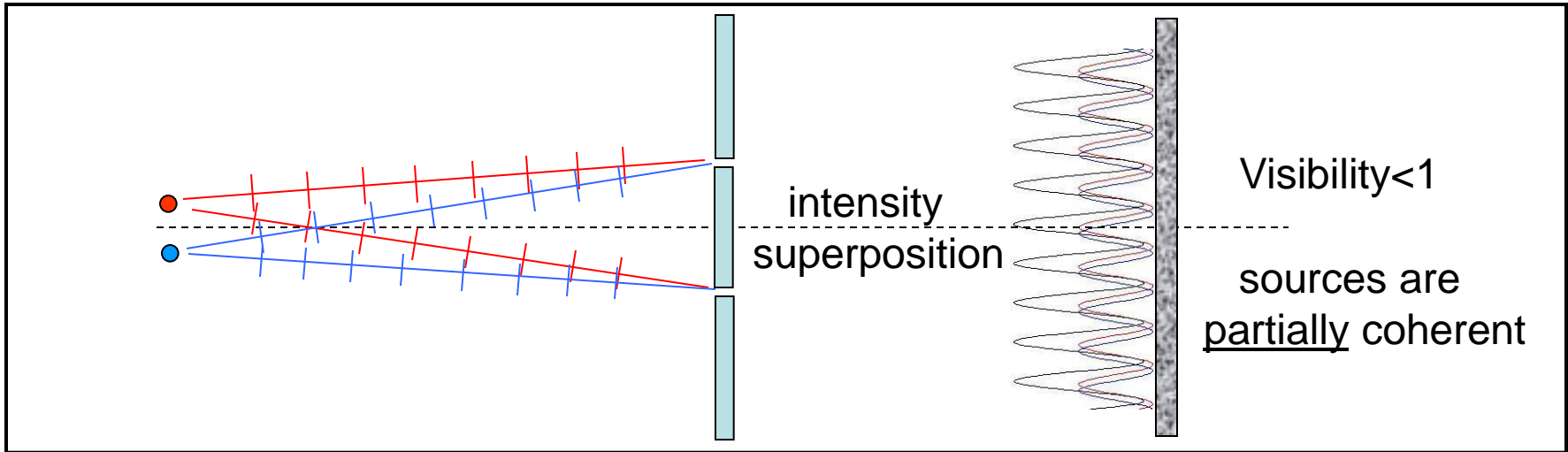
Γ_{12} is the degree of mutual-coherence

if $r < r_0$ (correlation length), then waves are coherent
and light interferes

if $r > r_0$ then coherence is lost
and light waves do not interfere

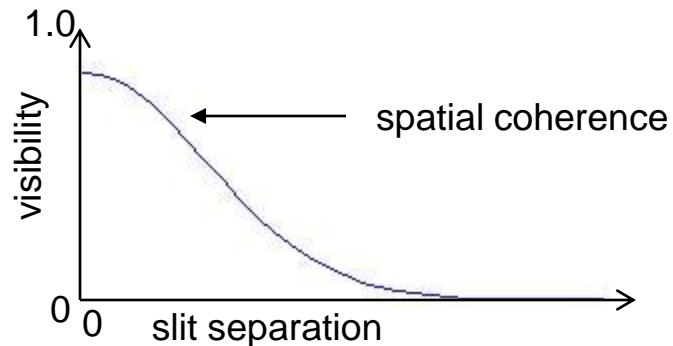


Spatial Coherence - Two Sources



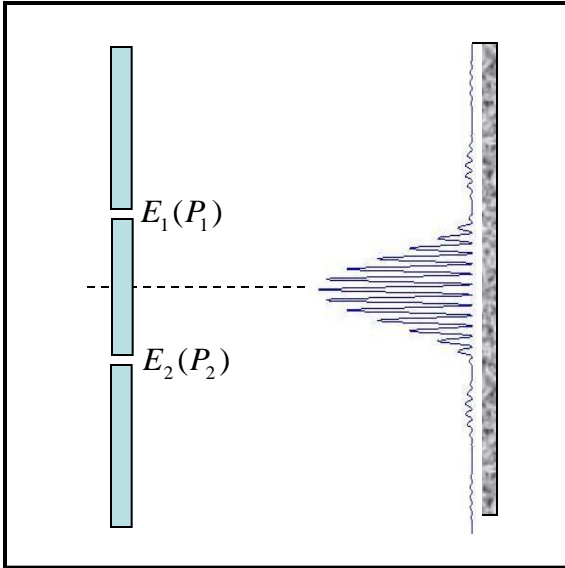
As slits separate

- 1) 'cosine' frequency goes up
- 2) superposition decoheres
- 3) visibility decreases



*preview: visibility is Fourier Transform
of source intensity*

Mutual Coherence of field at two points



Solve for intensity on the screen

$$I(\theta) = E_T^* E_T = (E_1 + E_2 e^{+i\theta})^* (E_1 + E_2 e^{-i\theta})$$

$$I(\theta) = E_1^2 + E_2^2 + E_1^* E_2 e^{-i\theta} + E_1 E_2^* e^{+i\theta}$$

Now take the time average

$$I(\theta) = E_{01}^2 + E_{02}^2 + E_{01} E_{02} (|\gamma| e^{-i\gamma} e^{-i\theta} + |\gamma| e^{+i\gamma} e^{+i\theta})$$

$$I(\theta) = E_{01}^2 + E_{02}^2 + E_{01} E_{02} |\gamma| \cos(\gamma + \theta)$$

$$Visibility = \frac{I_{Max} - I_{Min}}{I_{Max} + I_{Min}} = \frac{2E_{01}E_{02}|\gamma(P_1, P_2)|}{E_{01}^2 + E_{02}^2}$$

For $E_{01} = E_{02}$,

$$Visibility = \frac{I_{Max} - I_{Min}}{I_{Max} + I_{Min}} = |\gamma(P_1, P_2)|$$



Mutual Coherence (cont'd)

For the case when $I_1=I_2$, we have

$$Visibility = \frac{I_{Max} - I_{Min}}{I_{Max} + I_{Min}} = |\gamma(P_1, P_2)|$$

Where $\gamma = |\gamma(P_1, P_2)|e^{i\gamma} = \langle E_1 * E_2 \rangle$

has various descriptive labels

“mutual intensity”

“mutual coherence function”

“complex degree of coherence”

“correlation function”

“fringe parameter”

...fringe parameter measurement is central to all problems involving coherence



Putting it all together...

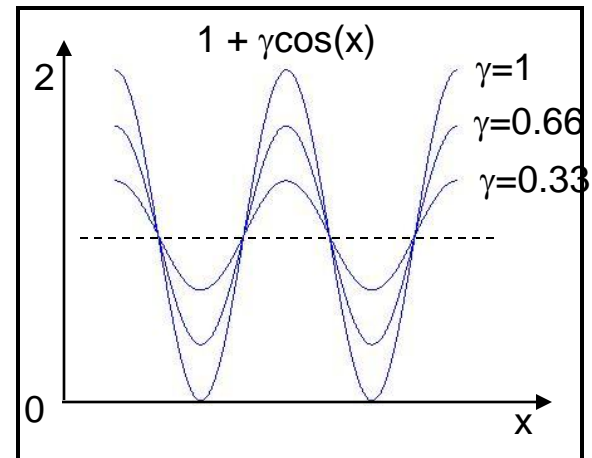
$$I(\theta) = E_{01}^2 + E_{02}^2 + E_{01}E_{02}|\gamma|\cos(\gamma + \theta)$$

$$I(y) = I_0 \underbrace{\left[\sin c \left(\frac{2\pi a}{\lambda R} y \right) \right]^2}_{\text{Single-Slit}} \cdot \underbrace{\left[1 + |\gamma| \cos \left(\frac{2\pi d}{\lambda R} y + \Phi \right) \right]}_{\text{Two-Slit}}$$

beam intensity
(equal both slits)

visibility factor (mutual coherence)

$$\text{Visibility} = \frac{I_{Max} - I_{Min}}{I_{Max} + I_{Min}} = |\gamma|$$



Alan will derive in-depth with chromatic effects



Van-Cittert/Zernike Theorem

“Visibility is the Fourier transform of source intensity”

$$\gamma(v) = \int I(y) e^{i2\pi v y} dy \quad I(y) = \text{intensity distribution}$$

$$v = \frac{d}{\lambda L} = \text{spatial - frequency}$$

$d = \text{slit width}$
 $\lambda = \text{wavelength}$
 $L = \text{source to slits}$

In two dimensions: $\gamma(v_x, v_y) = \iint I(x, y) e^{i2\pi(v_x x + v_y y)} dx dy$

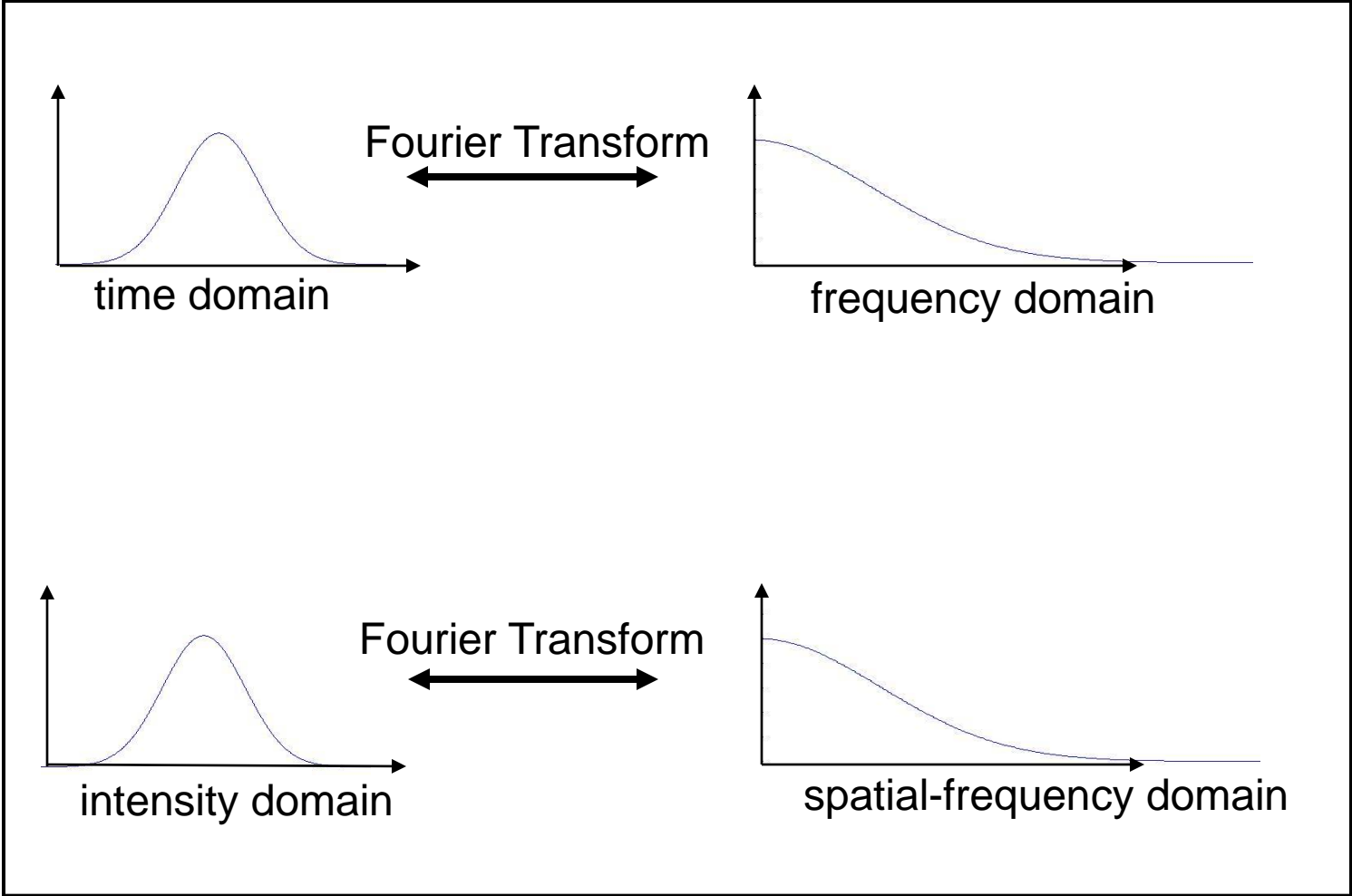
For a Gaussian, thermal-light source distribution

$$\gamma(d) = e^{-\frac{d^2}{2\sigma_d^2}} \quad (\text{one dimension})$$

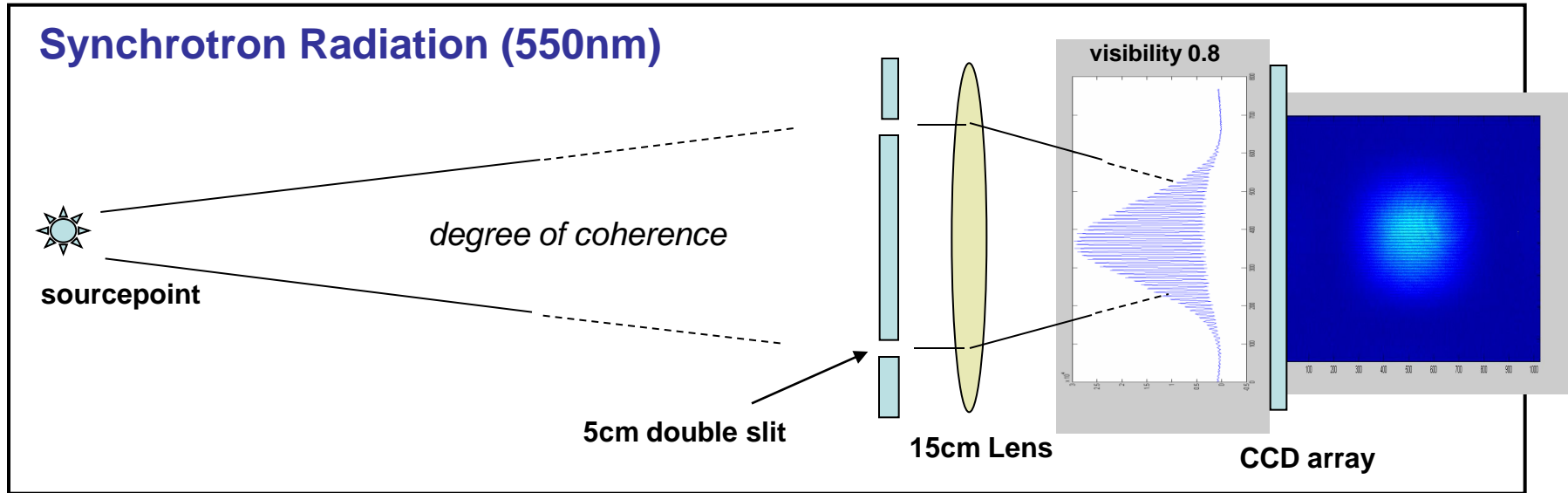
where $\sigma_d = \frac{2\pi\sigma_y}{\lambda L} = \text{spatial frequency characteristic}$



Fourier Transform Pairs



Interferometric Beam Size Measurement



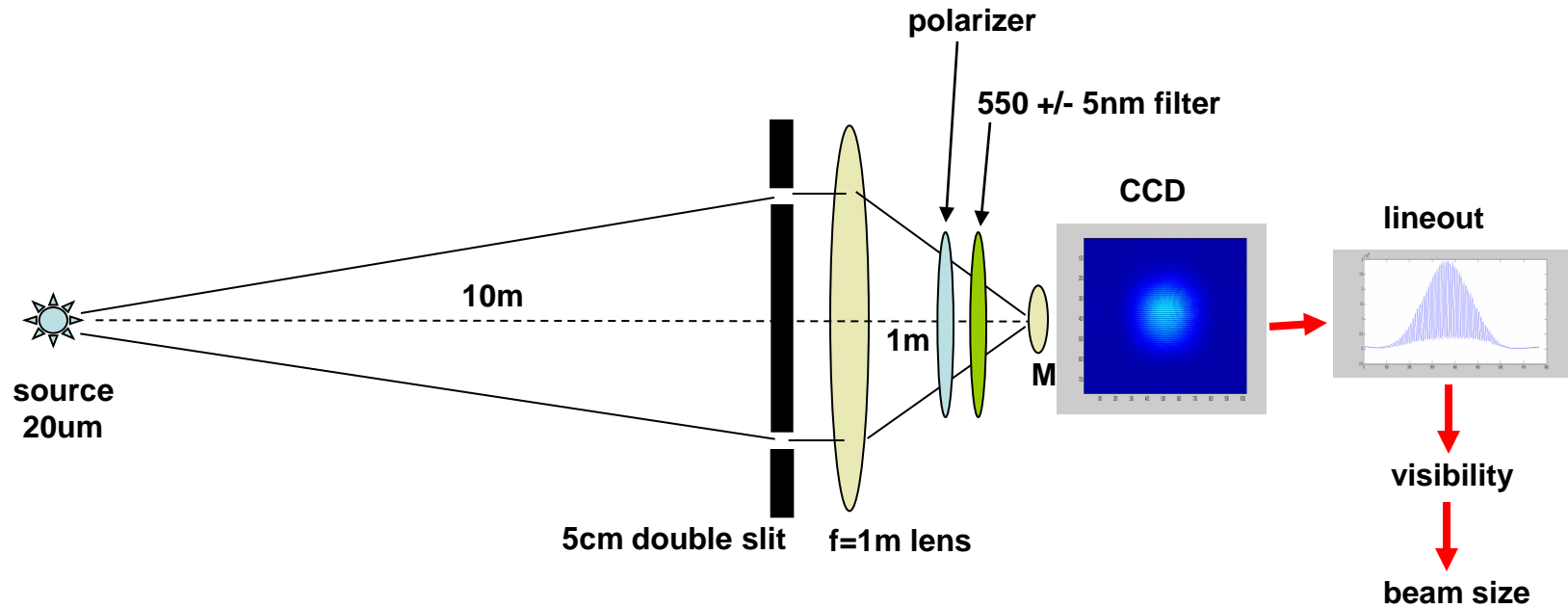
For a Gaussian source

$$\gamma(d) = e^{\frac{-d^2}{2\sigma_d^2}} \quad \sigma_d = \frac{\lambda L}{2\pi\sigma_y} = \text{spatial frequency characteristic}$$

- 1) Measure $\gamma(d)$ [visibility as a function of slit separation]
- 2) Solve for characteristic width σ_d
- 3) Infer beam size from: $\sigma_y = \lambda L / 2\pi\sigma_d$



Typical Stellar Interferometer for SR Measurements





Typical System Parameters

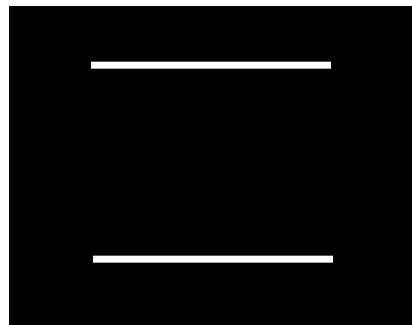
Source size: $\sigma_y = 20 \mu\text{m}$

Source-slit: $L = 10 \text{m}$

wavelength: $\lambda = 550 \text{nm}$

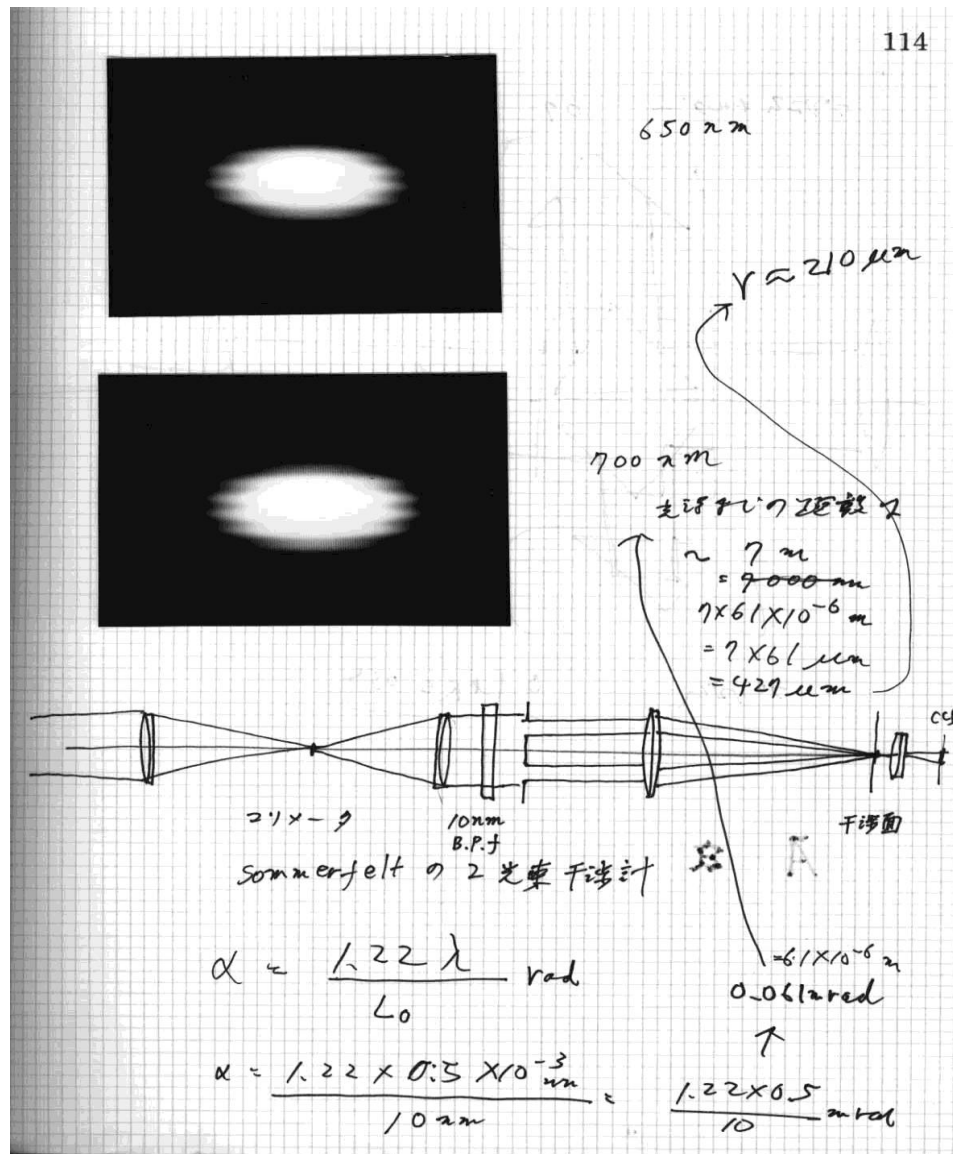
Visibility $\gamma(d) = e^{\frac{-d^2}{2\sigma_d^2}}$

$$\sigma_d = \frac{\lambda L}{2\pi\sigma_y} = \frac{550 \times 10^{-9} \cdot 10}{2\pi \cdot 20 \times 10^{-6}} = 44 \text{mm}$$



$a \sim 2 \text{ mm}$

$d \sim 40 \text{ mm}$

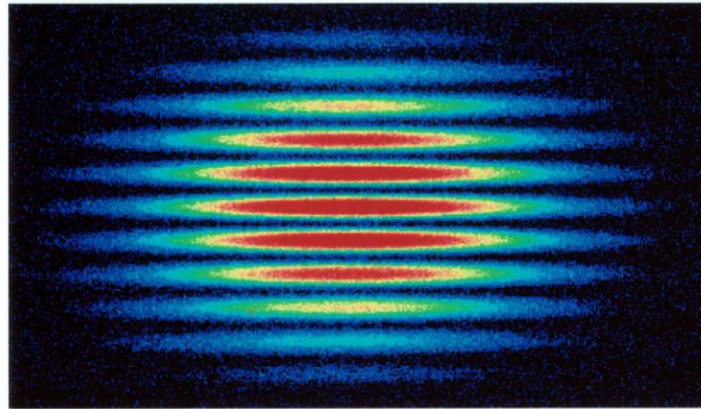


Result of beam size is $210 \mu\text{m}$

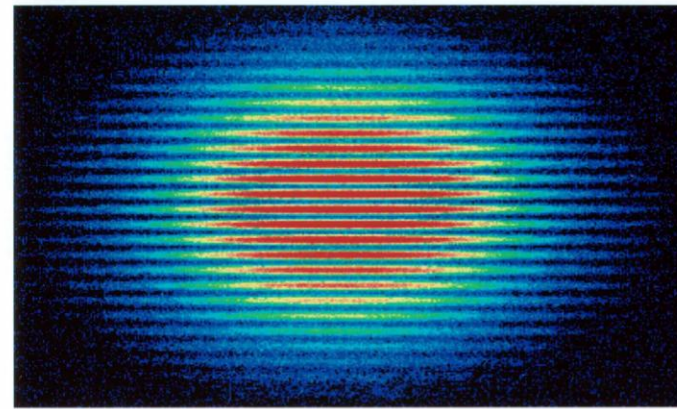


Vertical beam size at the SR center of Ritsumeikan university AURORA.

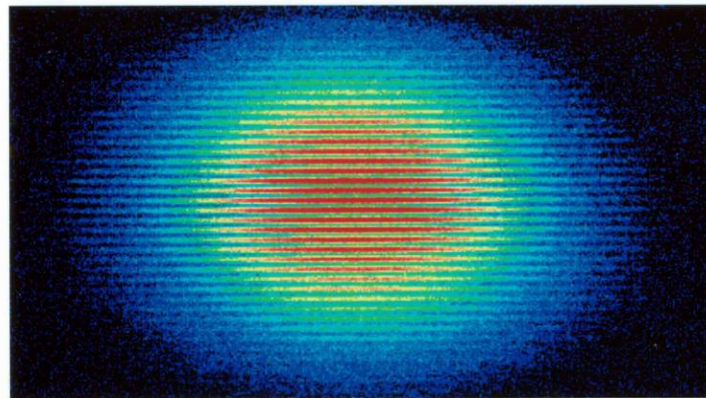
$\lambda = 550\text{nm}$



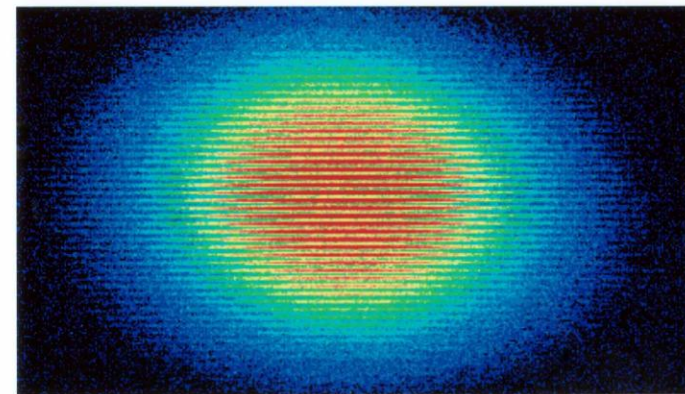
D=6.7mm (1.79mrad)



D=14.7mm (3.92mrad)



D=22.7mm (6.05mrad)

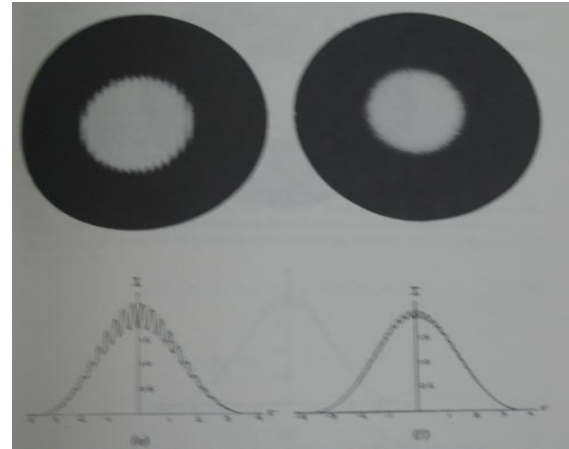
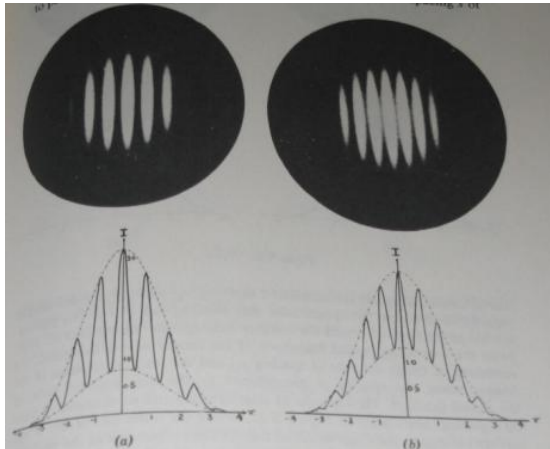


D=28.7mm (7.65mrad)

compliments: T. Mitsuhashi



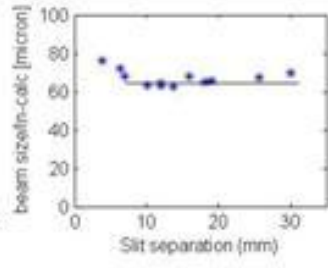
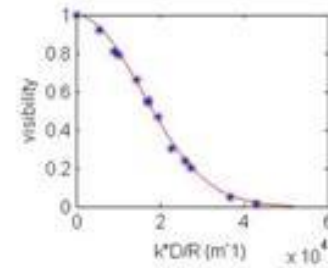
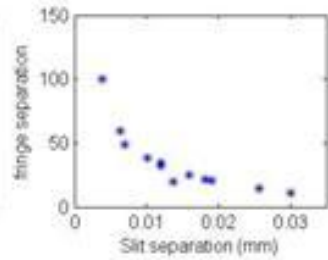
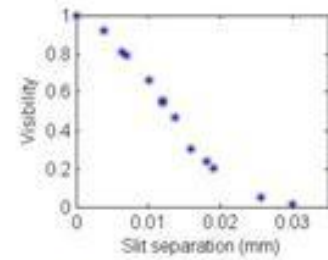
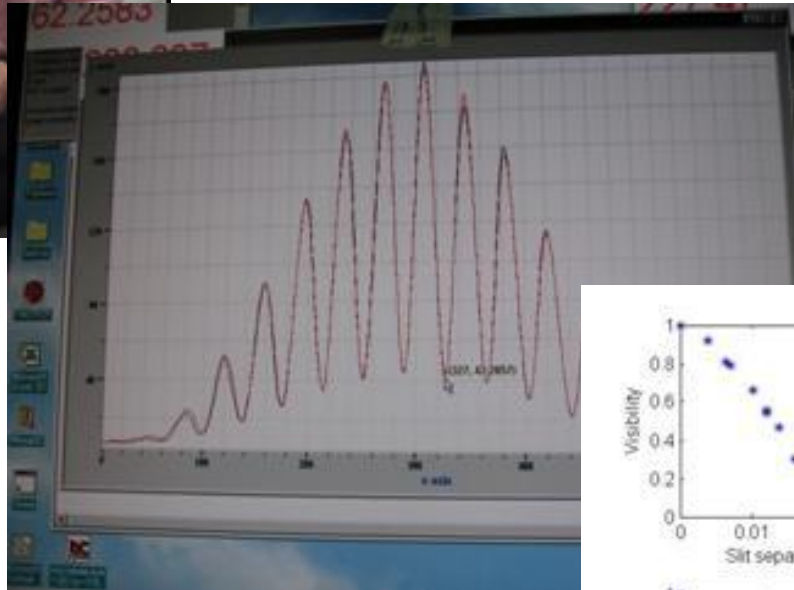
Example from Goodman/Statistical Optics



Note effect of vibration!



Photon Factory Laboratory





Beam Size Measurement (cont'd)

From a single measurement at slit separation d_o

$$\gamma(d_o) = e^{\frac{-d_o^2}{2\sigma_d^2}} \quad \text{solve for } \sigma_d = \sqrt{\frac{d_o}{\ln\left(\frac{1}{\gamma}\right)}}$$

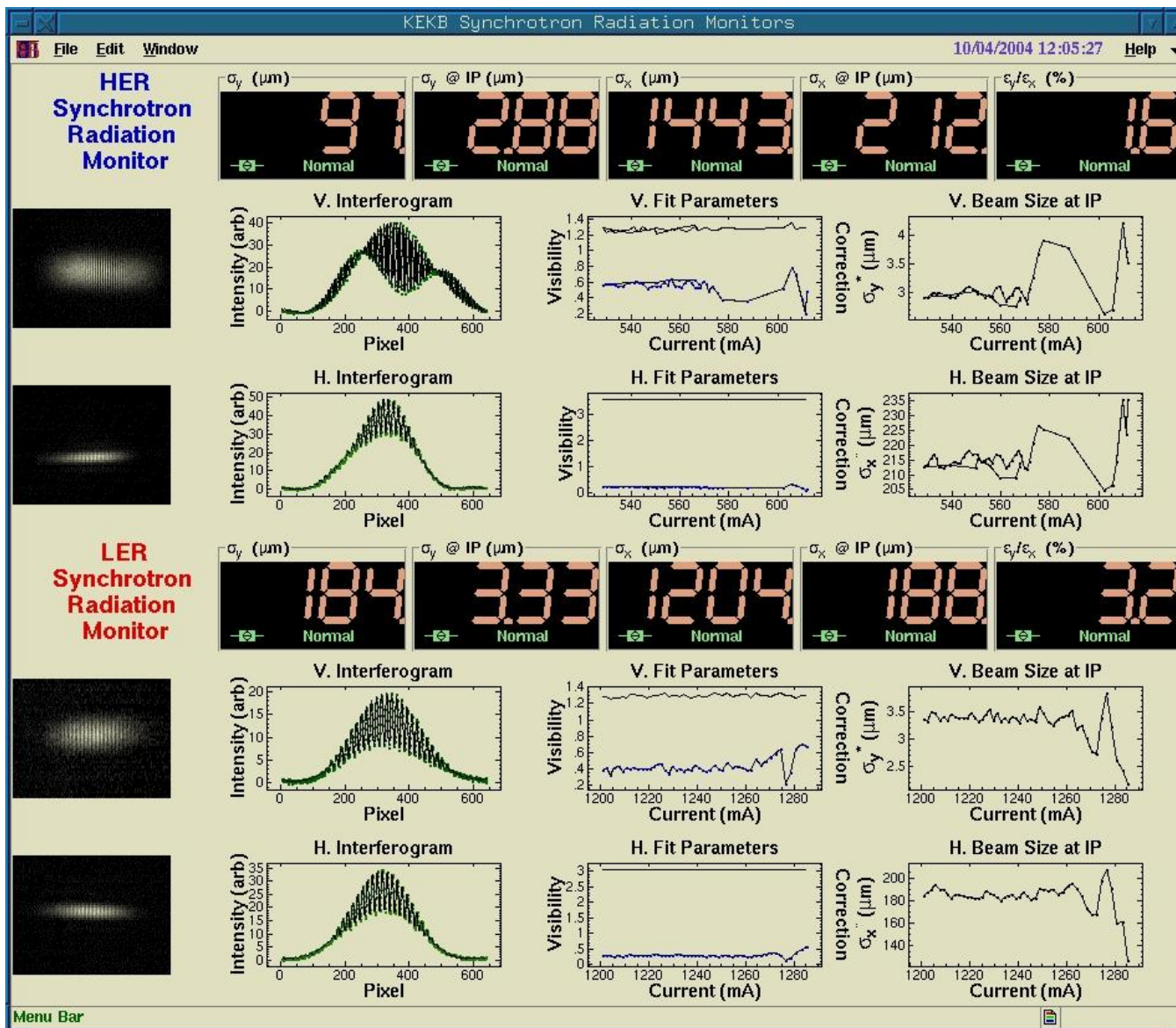
From $\sigma_y = \frac{2\pi}{\lambda L \sigma_d}$

$$\sigma_y = \frac{\lambda L}{\pi d_o} \sqrt{\frac{\ln\left(\frac{1}{\gamma}\right)}{2}}$$

on-line diagnostic for slit separation d_o



KEK-B system





PEP-II system

Synchrotron Light

Synchrotron Light Software Instr...

LER Interferometer

LER Fringe Calculation

LER Fringe Strip Chart

HER Interferometer

HER Fringe Calculation

HER Fringe Strip Chart

LER X-Ray Image rotation

Start 18 LabVIEW Develop... 07:52

LER Interferometer Data:

Whole	Region
Max 663	559
Mean 37	145
Min 0	5
Centroid X 1.12E+48	
Centroid Y -30.5	-1

LER Fringe Calculation Data:

Good FR?	Estimate	Fit Result	Error Bar
Baseline	0.035	0.034	0.001
Amplitude, Slit 1	1	0.304	0.001
Amplitude, Slit 2	1	0.304	0.001
Angle on Slits (urad)	17.41	-3.37	0.82
Phase Offset (deg)	0.00	-42.70	2.21
Outer Jaw (mm)	1.90	1.89	0.00
Inner Jaw (mm)	1.38	1.46	0.00
Beam Size (mm RMS)	0.180	0.180	0.001
Visibility	0.56	0.57	
Camera z Position (mm)	-20.0		
Chi Square & Error Code		0.93	0

HER Interferometer Data:

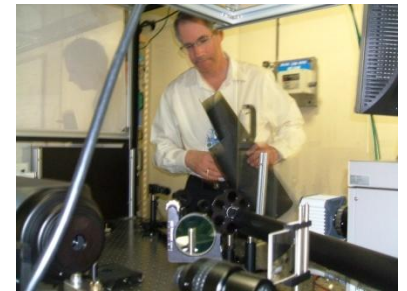
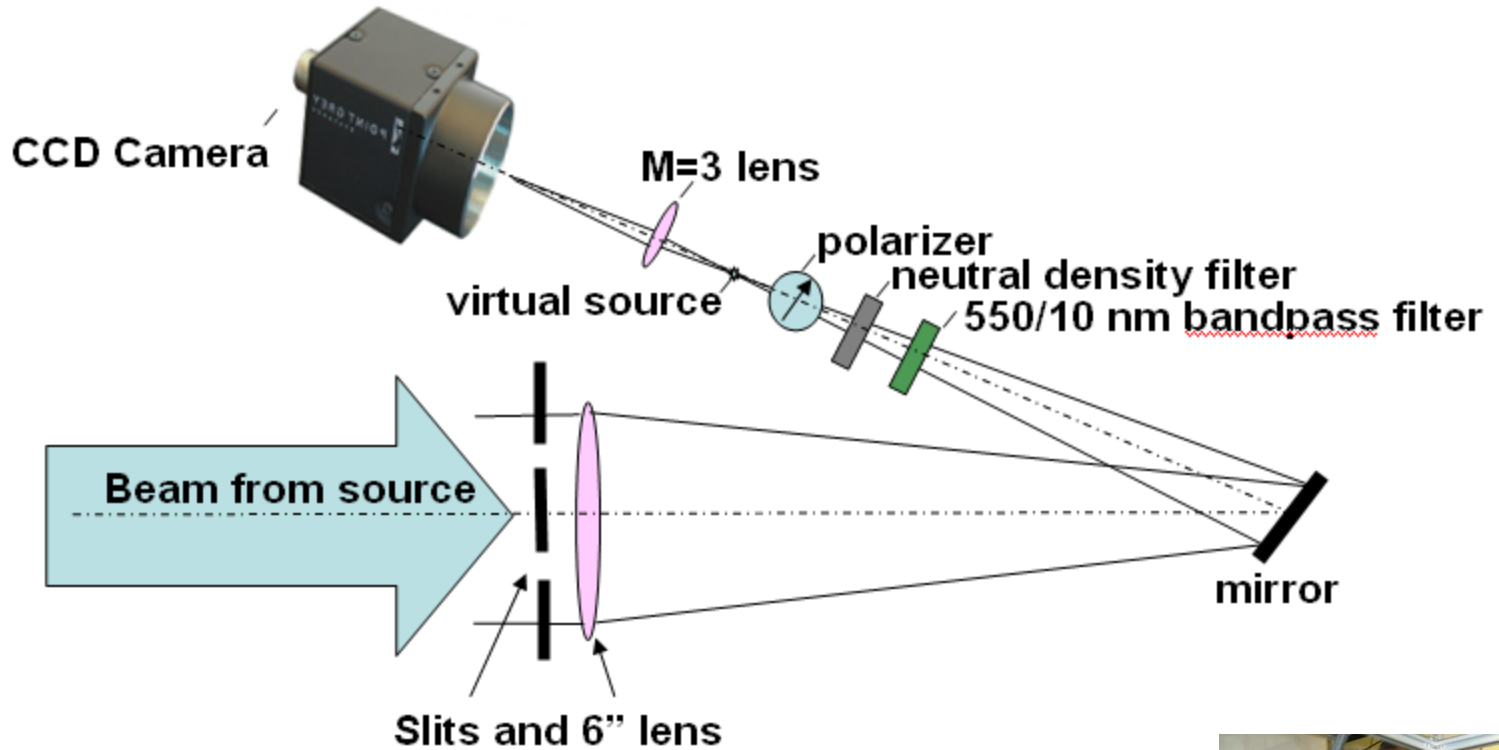
Whole	Region
Max 676	431
Mean 41	104
Min 4	10
Centroid X 1.25E+1	
Centroid Y 4.1	43

HER Fringe Calculation Data:

Good FR?	Estimate	Fit Result	Error Bar
Baseline	0.044	0.047	0.001
Amplitude, Slit 1	1	0.272	0.001
Amplitude, Slit 2	1	0.272	0.001
Angle on Slits (urad)	-174.47	-49.74	1.19
Phase Offset (deg)	0.00	64.79	3.95
Outer Jaw (mm)	2.22	2.22	0.00
Inner Jaw (mm)	1.90	1.91	0.00
Beam Size (mm RMS)	0.171	0.169	0.001
Visibility	0.60	0.54	
Camera z Position (mm)	-65.0		
Chi Square & Error Code		3.79	0

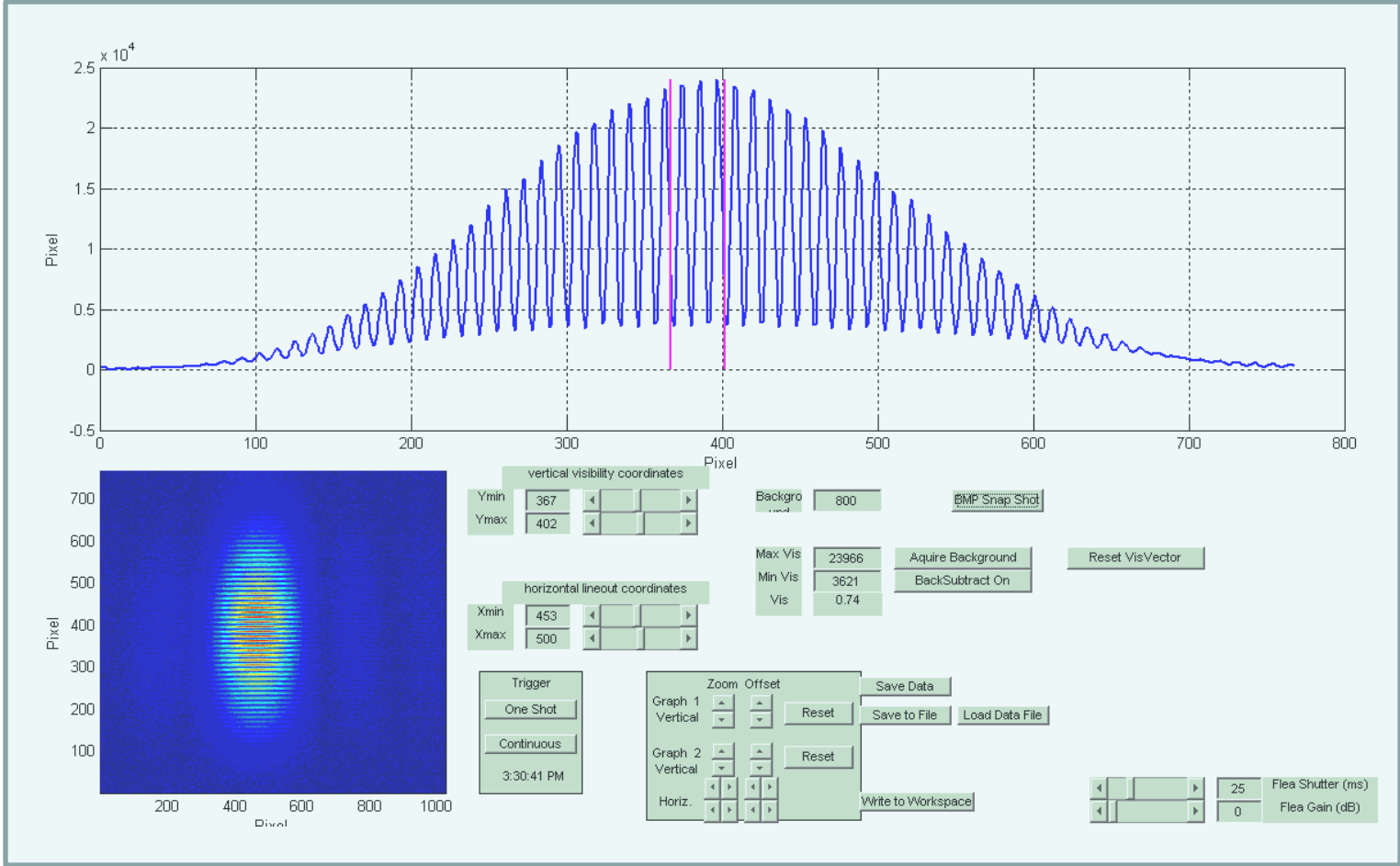


SPEAR3 system



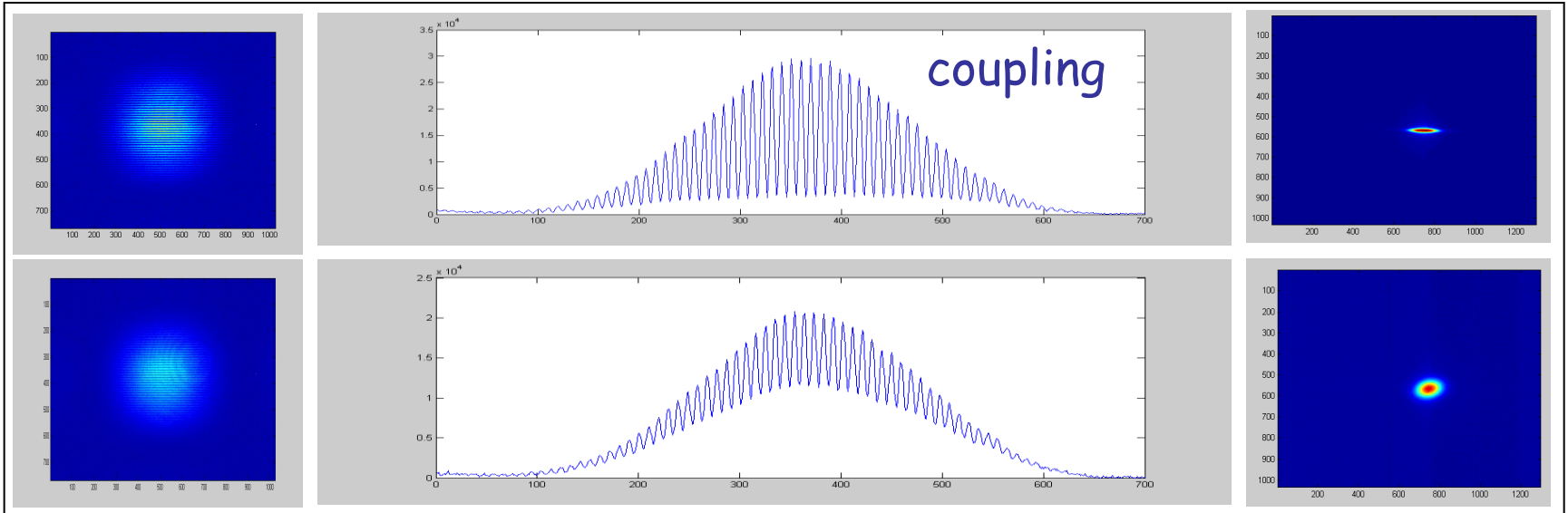
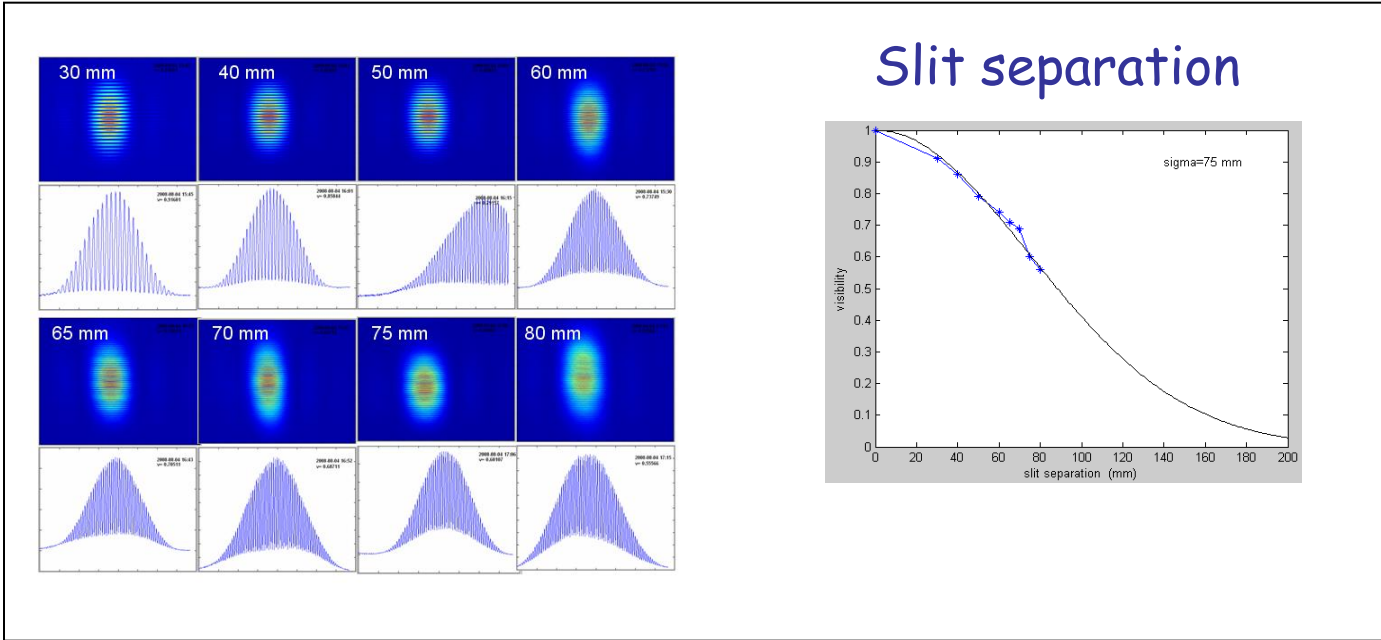


SPEAR3 system (cont'd)



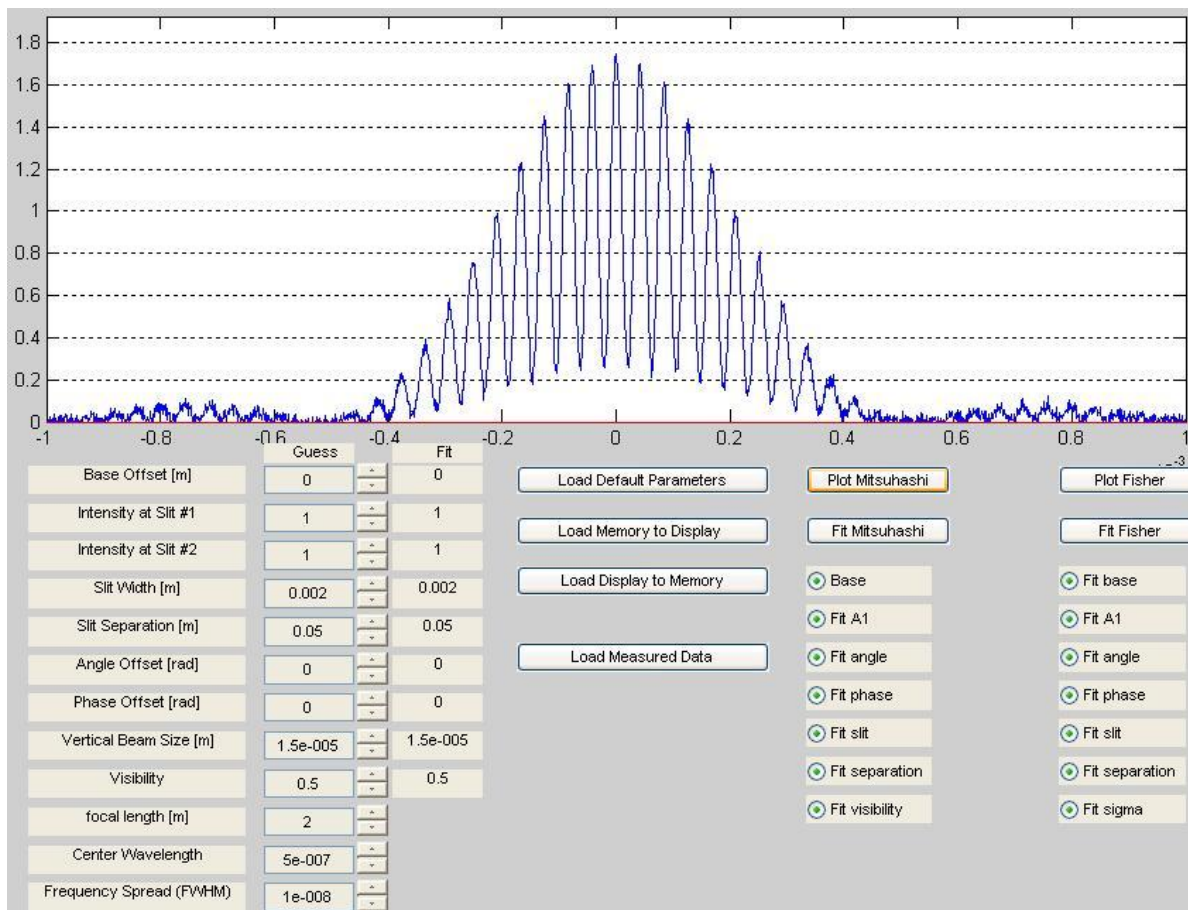


SPEAR-3 measurements





USPAS Simulator





Practical Issues

- Thermal distortion of mirrors – wavefront distortion
- Precision control of slit width (I_1 and I_2)
- Depth of field effects
- CCD camera linearity
- Table vibrations
- Readout noise
- Beam stability
- Numerical fitting



Michelson's Interferometer - Summary

- Interferometers useful below the diffraction limit
- Two-slit Interference – Young's experiment
- Diffraction from a single slit
- Extended Source – Partial Coherence
- Visibility and the Mutual Coherence function
- Van-Cittert/Ziernike theorem: Fourier XFRM
- Stellar Interferometers for SR applications