



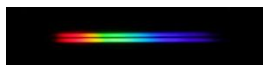
# Beam Diagnostics with Synchrotron Radiation - An Overview

J. Corbett, W. Cheng, A. Fisher and W. Mok  
US Particle Accelerator School  
January 18-22, 2010

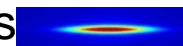
- Motivation

- Electron beam properties

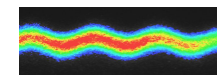
- SR beam properties



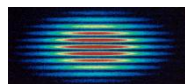
- Transverse imaging – cameras, pinholes, Fresnel lenses



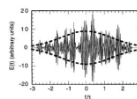
- Fast imaging – gated ICCD and streak cameras



- Stellar Interferometer



- Fluctuation measurements



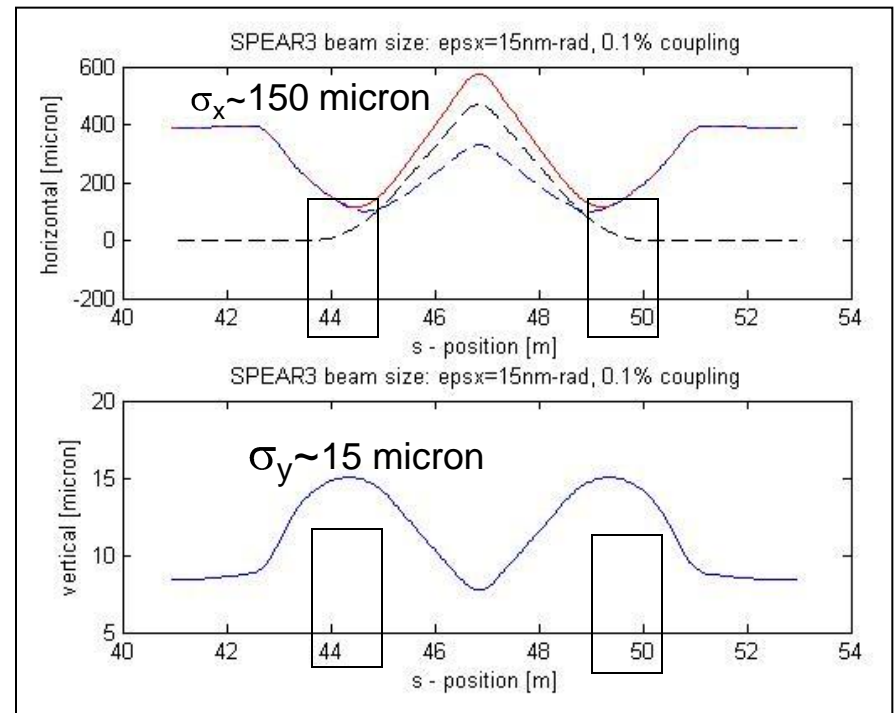
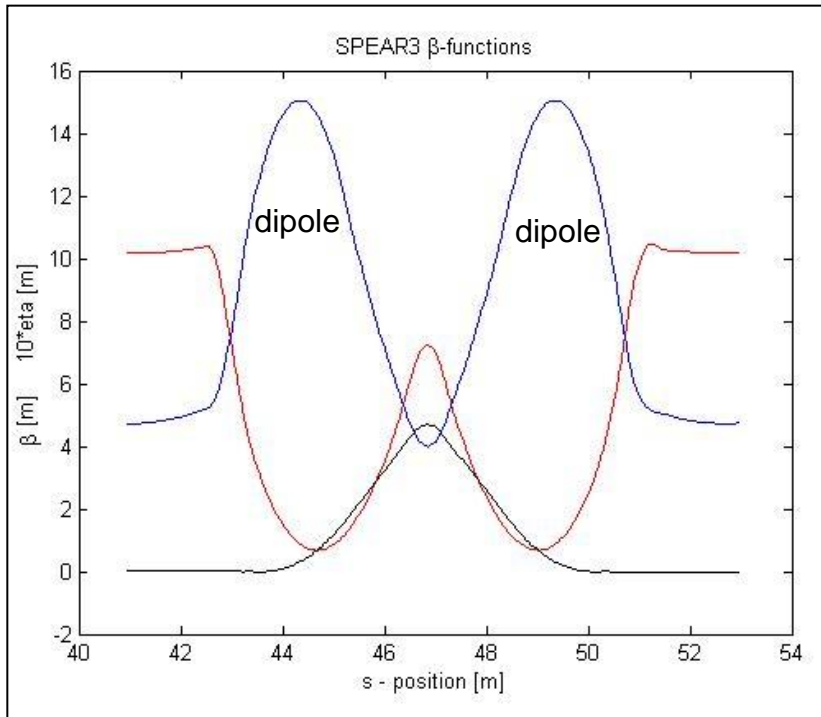


## Motivations for SR diagnostics

- 'eye' into the accelerator
- faithful photon image reproduces electron beam distribution (x,y,z)
- optics verification, coupling, brightness
- impedance and instabilities
- other techniques less accurate (e.g. scraper, RMA)



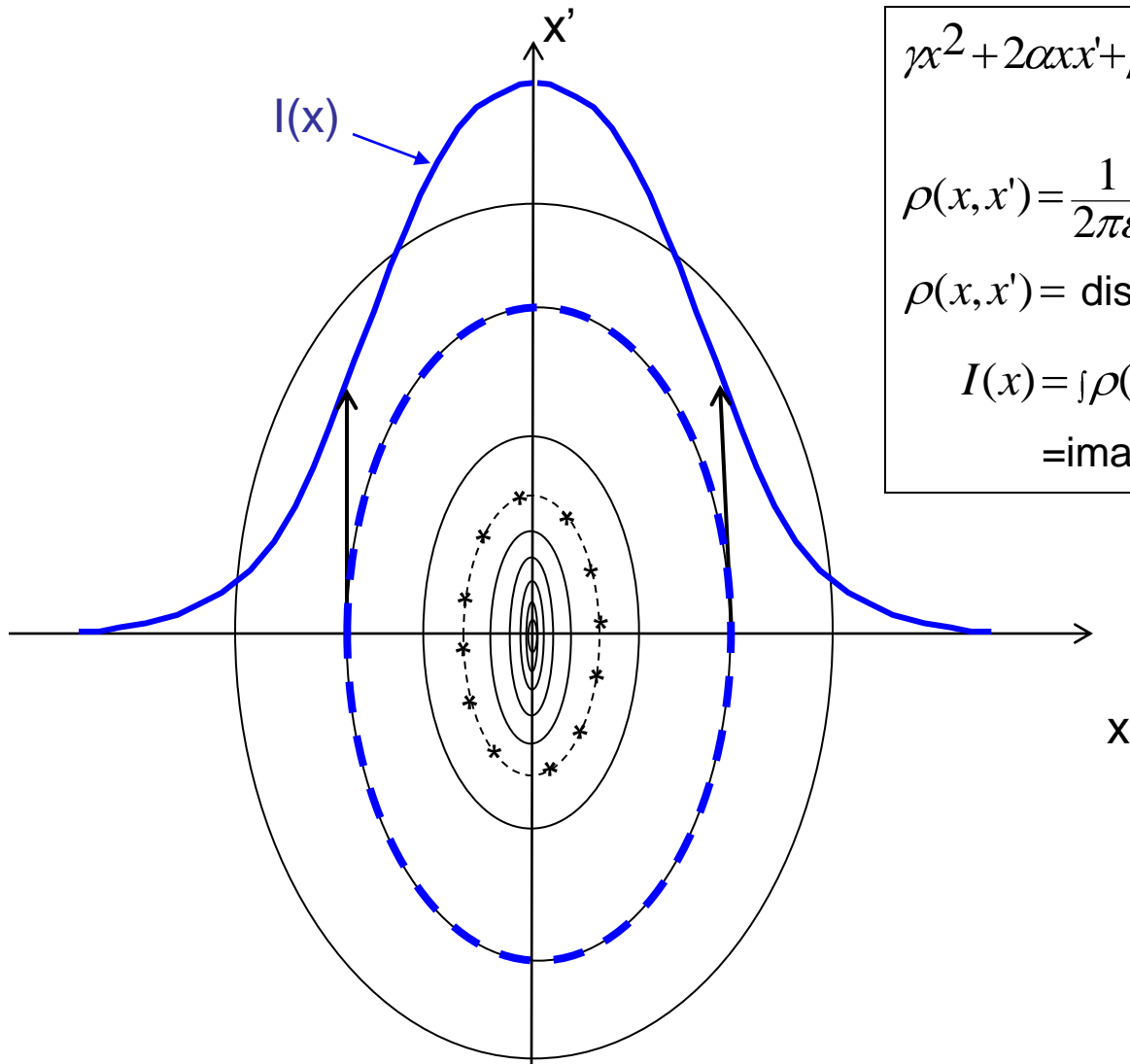
# Electron beam properties: $\beta$ -functions and beam size



Hofmann: chapter 13



# Electron beam properties: Phase Space



$\gamma x^2 + 2\alpha x x' + \beta x'^2 = \text{invariant oscillation amplitude}$

$$\rho(x, x') = \frac{1}{2\pi\epsilon} e^{-(\gamma x^2 + 2\alpha x x' + \beta x'^2)/2\epsilon}$$

$\rho(x, x') = \text{distribution in phase space}$

$$I(x) = \int \rho(x, x') dx' = A e^{-x^2/2\epsilon\beta}$$

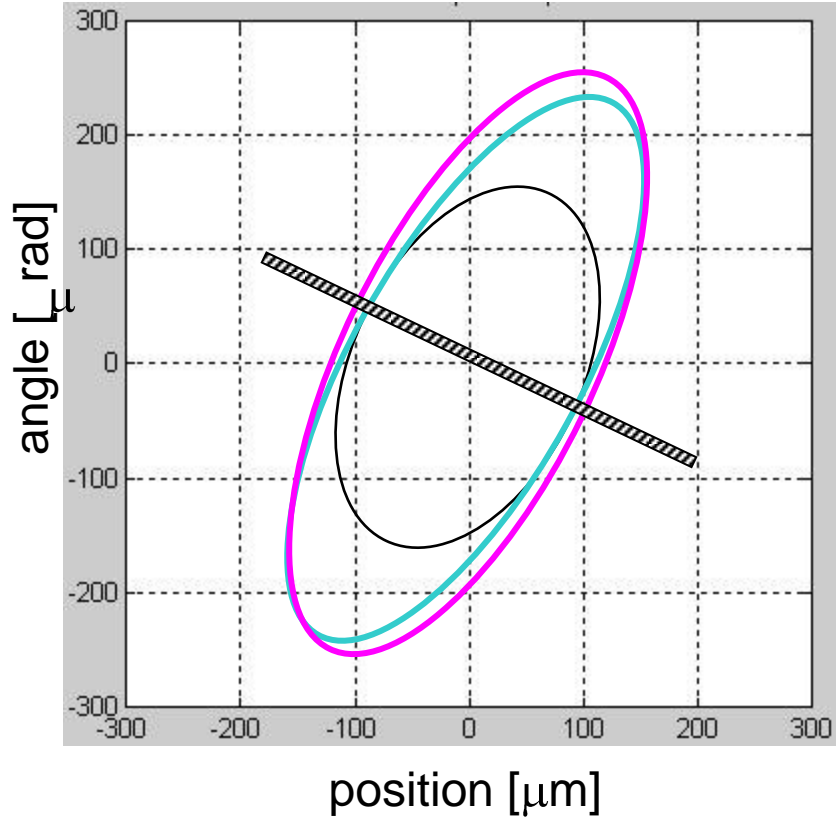
=image projected on screen

Hofmann: chapter 13



# Photon beam properties: Phase Space

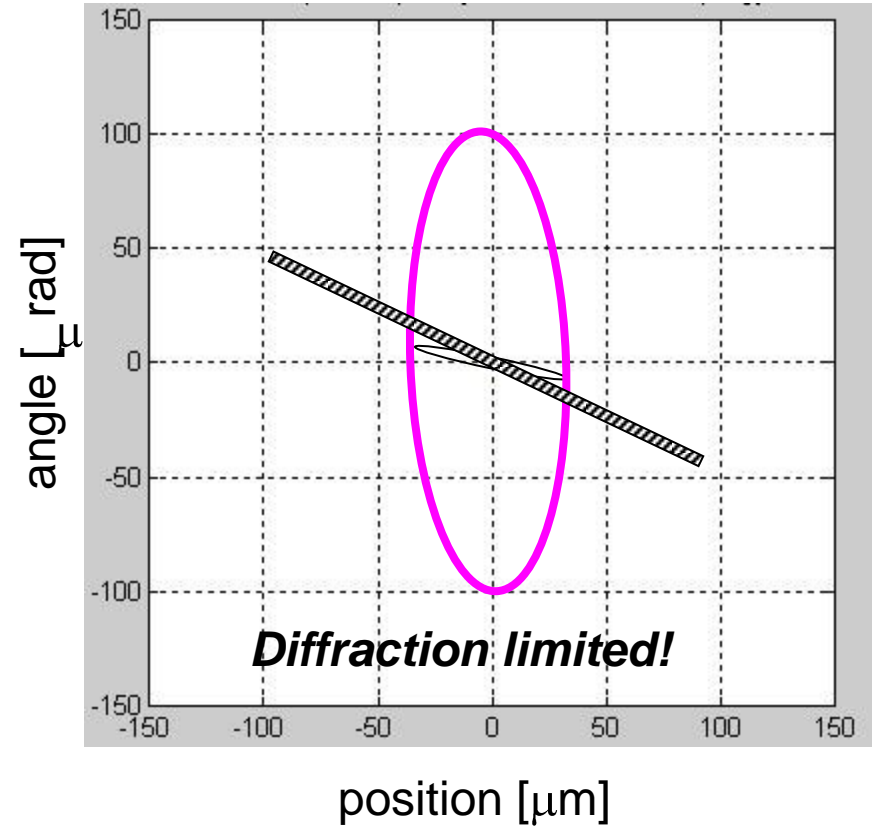
## SPEAR3: Horizontal



$$\sigma_x^2 = \varepsilon_x \beta_x + (\eta_x \delta)^2$$

$$\sigma_{x'}^2 = \varepsilon_x \gamma_x + (\eta'_x \delta)^2 + \sigma_r^2$$

## SPEAR3: Vertical



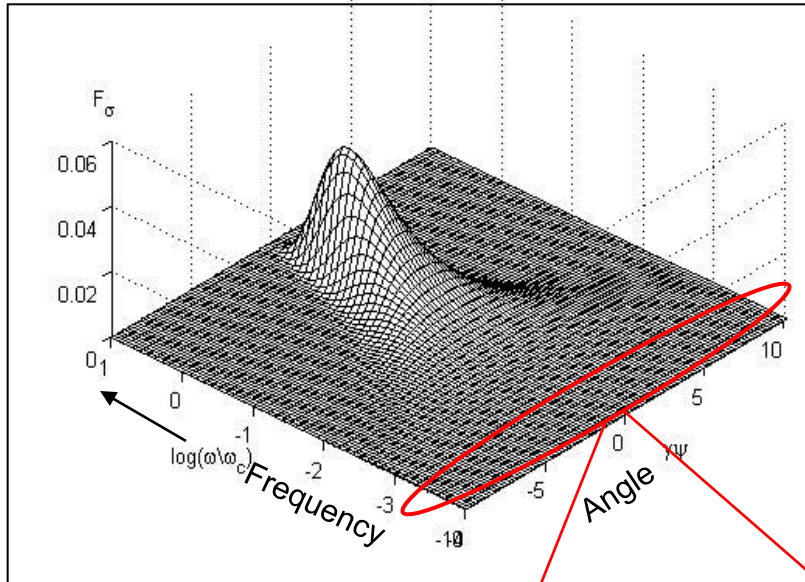
$$\sigma_y^2 = \varepsilon_y \beta_y + (\eta_y \delta)^2$$

$$\sigma_{y'}^2 = \varepsilon_y \gamma_y + (\eta'_y \delta)^2 + \sigma_r^2$$

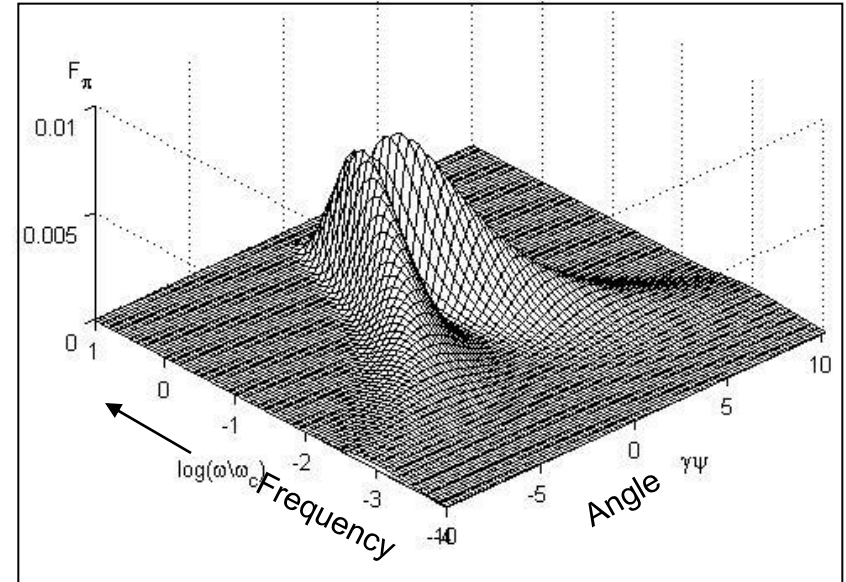


# Photon beam properties: Angular spectral power density

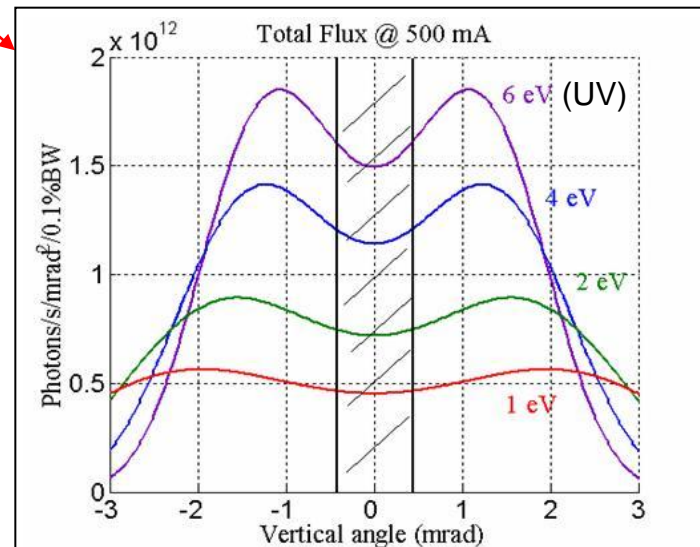
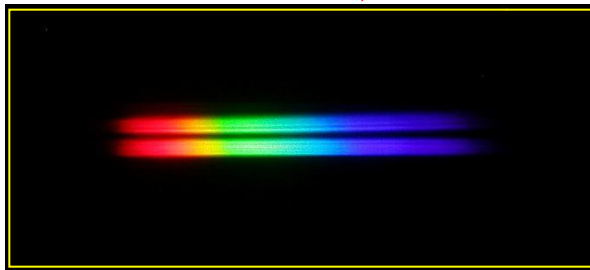
Horizontal Polarization



Vertical Polarization



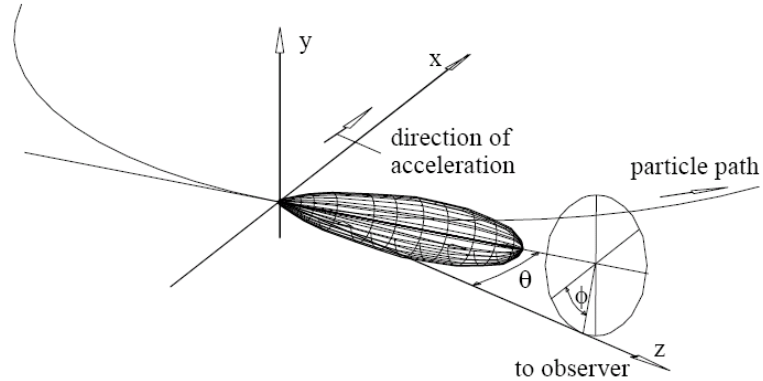
visible component



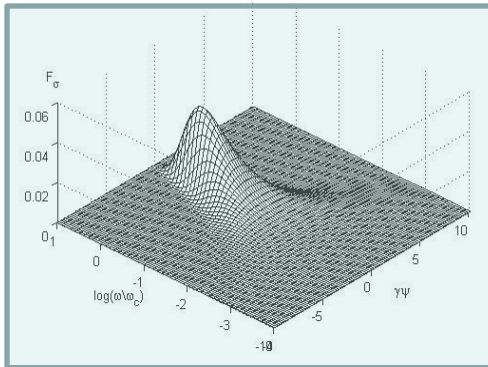
Hofmann: chapter 5



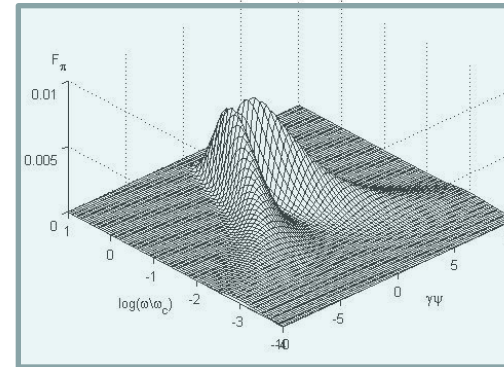
# Angular spectral power density (Schwinger, 1946)



Horizontal Polarization



Vertical Polarization



$$F_{s\sigma}(\theta, \psi) = \left(\frac{3}{2\pi}\right)^3 \left(\frac{\omega}{2\omega_c}\right)^2 \left(1 + \gamma^2 \psi^2\right)^{-2} K_{2/3}^2 \left(\frac{\omega}{2\omega_c} \left(1 + \gamma^2 \psi^2\right)^{3/2}\right)$$

$$F_{s\pi}(\theta, \psi) = \left(\frac{3}{2\pi}\right)^3 \left(\frac{\omega}{2\omega_c}\right)^2 \gamma^2 \psi^2 \left(1 + \gamma^2 \psi^2\right)^{-2} K_{1/3}^2 \left(\frac{\omega}{2\omega_c} \left(1 + \gamma^2 \psi^2\right)^{3/2}\right)$$



## Photon beam properties: Beam power

### Example:

3 GeV and 200 ma current

$P_{SR} \sim 200\text{kW}$  (total)

visible beam line (1.5eV)

25mm aperture at 5 m (5 mrad)

$P_{SR} \sim 150\text{W}$

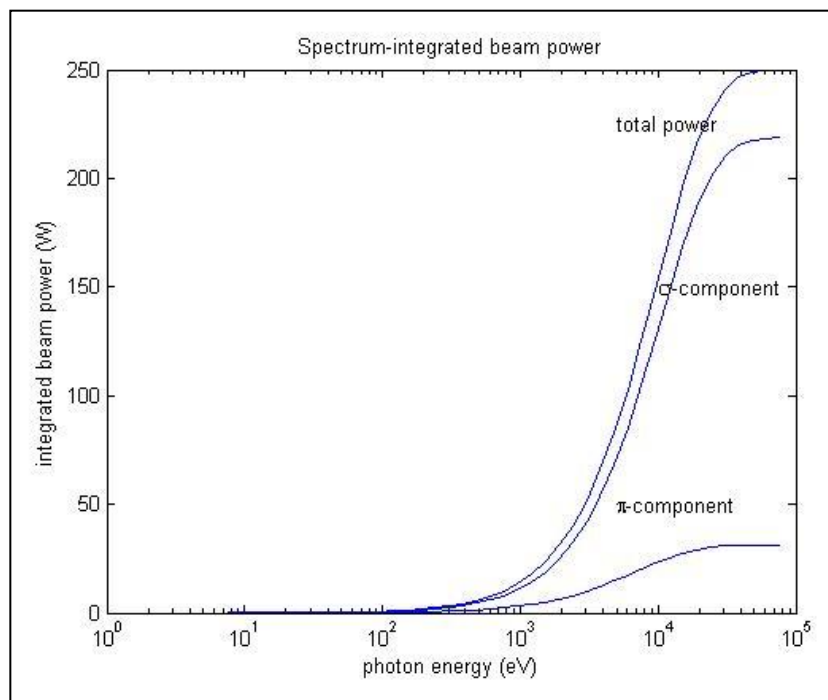
*lucky to get 100 $\mu\text{W}$  visible*

(Class I laser pointer)

pinhole camera (15keV)

25 $\mu\text{m}$  aperture at 5 m (5  $\mu\text{rad}$ )

$P_{SR} \sim 100\mu\text{W}$  before filter

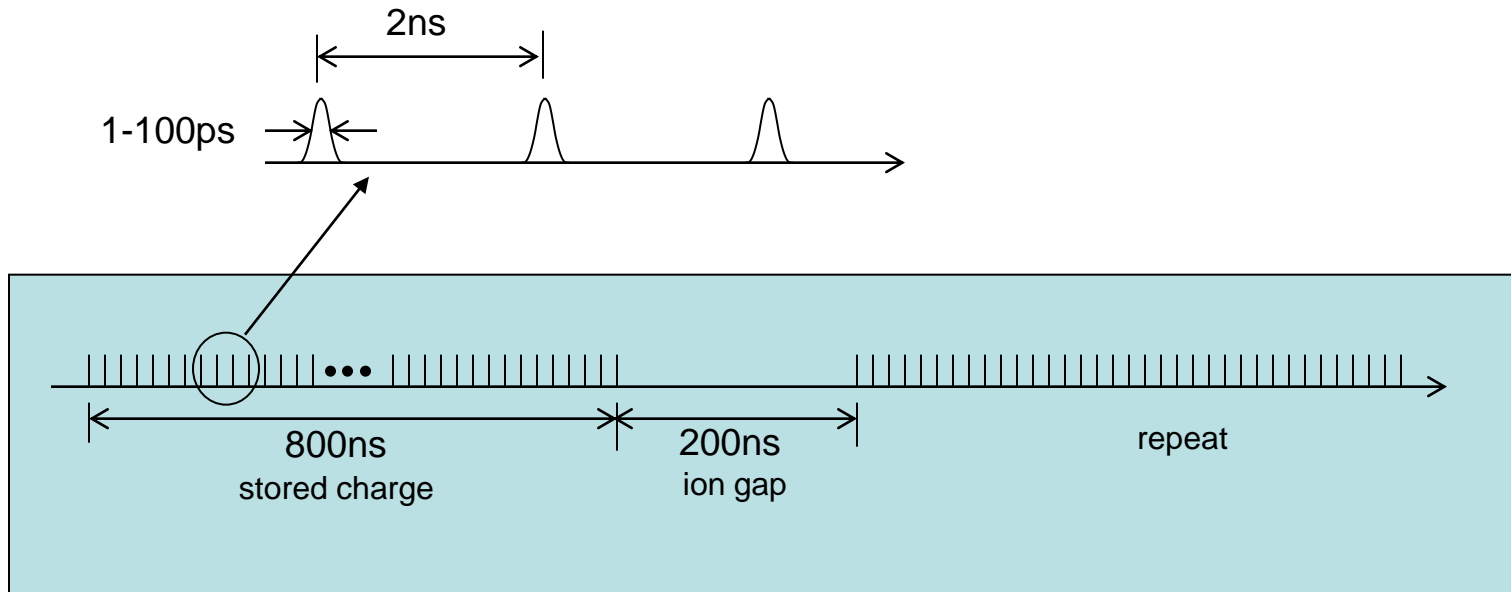


Hofmann: chapter 5



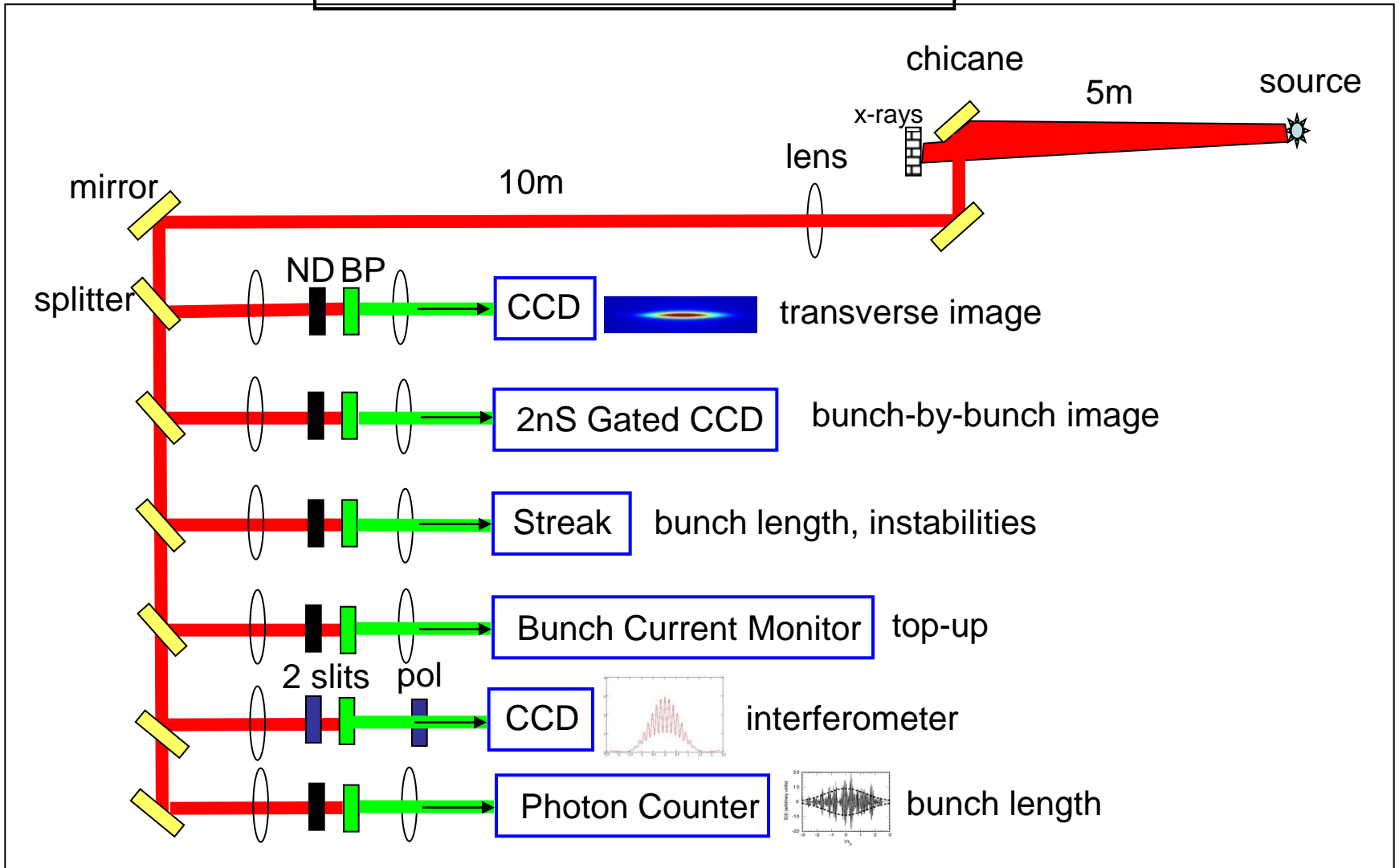


# Photon beam timing pattern - storage rings





# Visible beam line components





## Visible beam line components (cont'd)

### Beam line optics

Windows - quartz can pass down to about 220nm

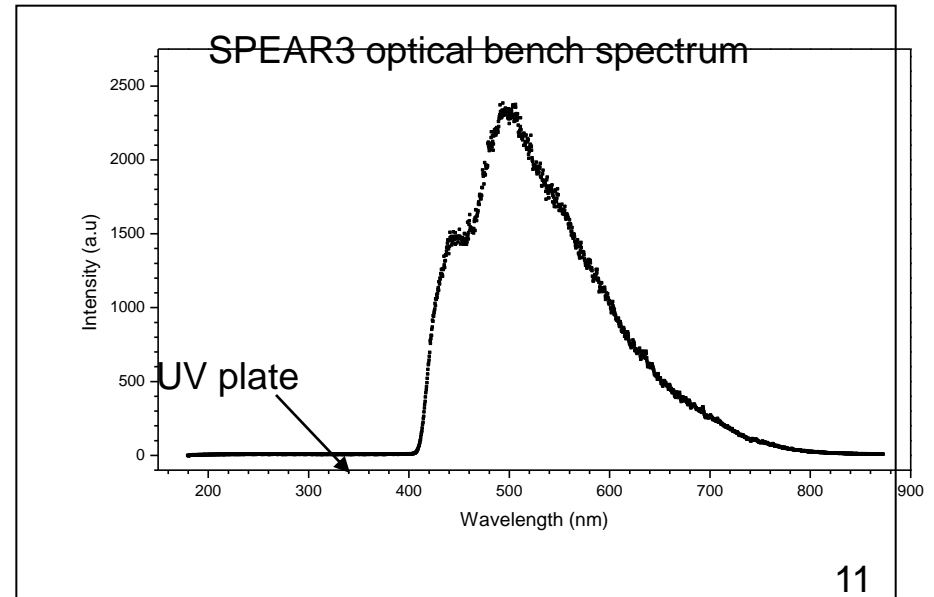
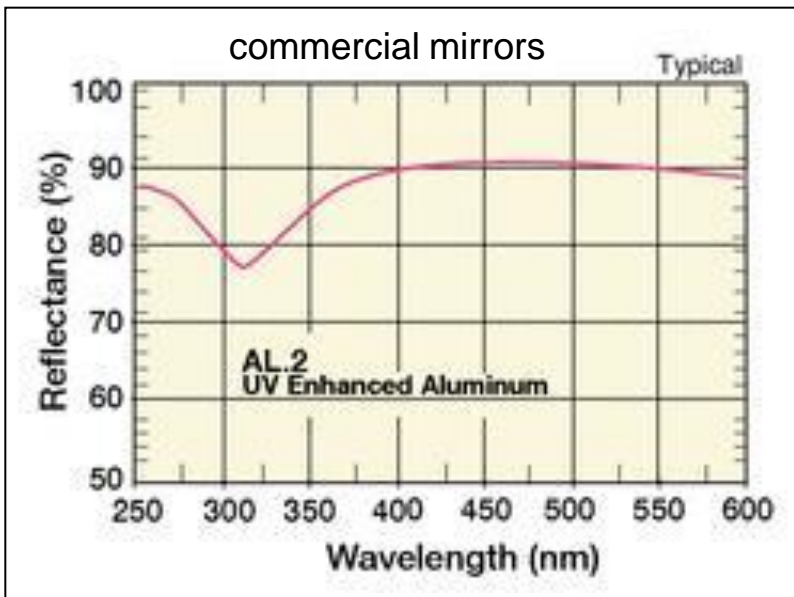
Mirrors – flat or focusing, UV enhanced

Lenses – focusing, defocusing, doublets, achromats >350nm

Filters –highpass and bandwidth to about 10nm FWHM

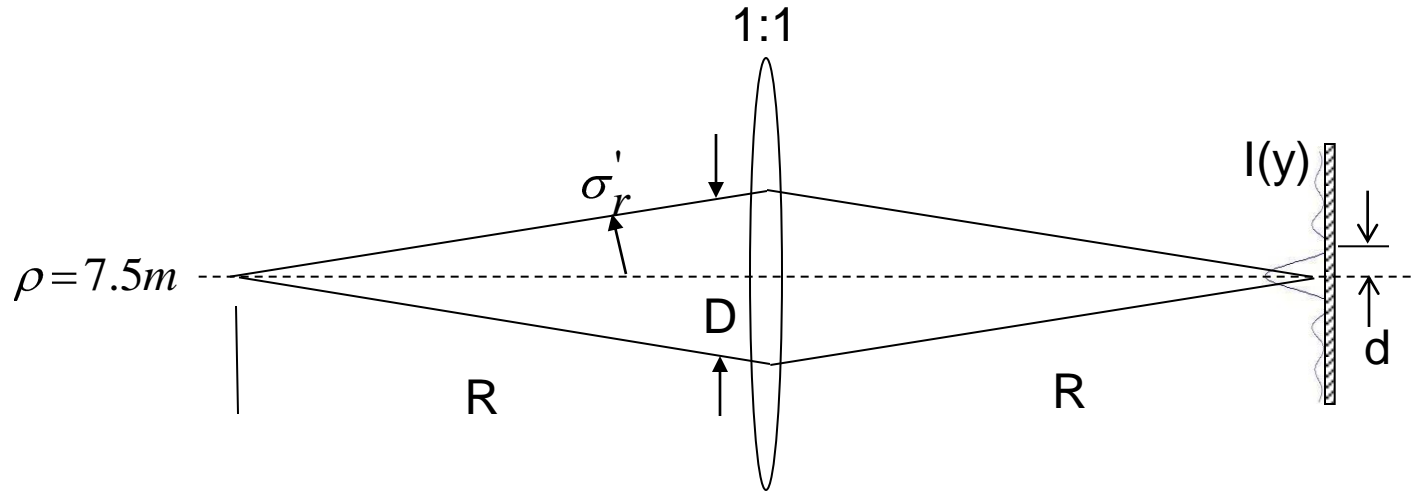
Slits and diaphragms – 1ms mechanical shutters, 10ps Pockel cells

About 90% transmission per element





# Diffraction Limited Resolution



For lens aperture:  $d \approx \frac{\lambda R}{2D}$       where for SR  $D = 2\sigma_r' R$

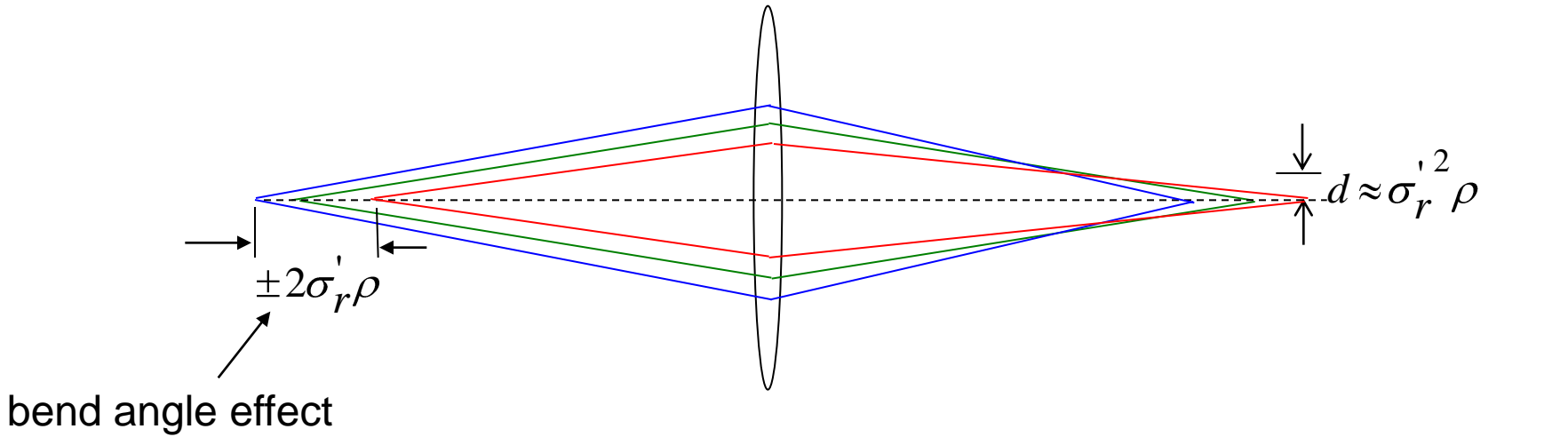
$d \approx \frac{1}{3} \cdot (\lambda^2 \rho)^{1/3}$        $\sigma_r' = 0.41 \cdot \lambda / \rho^{1/3}$       visible

$d \approx 40 \mu m$       at  $\lambda = 550 nm$

Hofmann: chapter 12

more diffraction later...

# Depth of Field



$$d \approx \sigma_r'^2 \rho$$

$$d \approx \frac{1}{3} \cdot (\lambda^2 \rho)^{1/3}$$

~same result as diffraction

(source length related to opening angle)

Hofmann: chapter 12



# Cameras Part-I: CCD's and Video

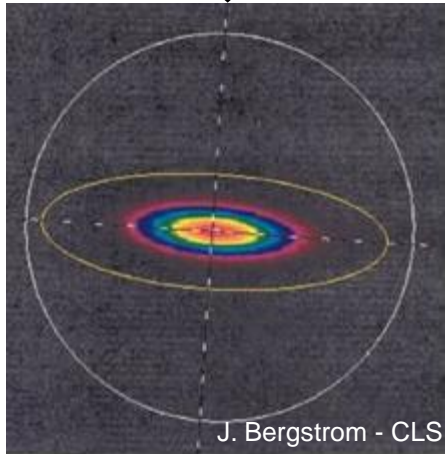


Frame Grabber

**Firewire®**  
Laser Beam Profilers  
From **Spiricon**



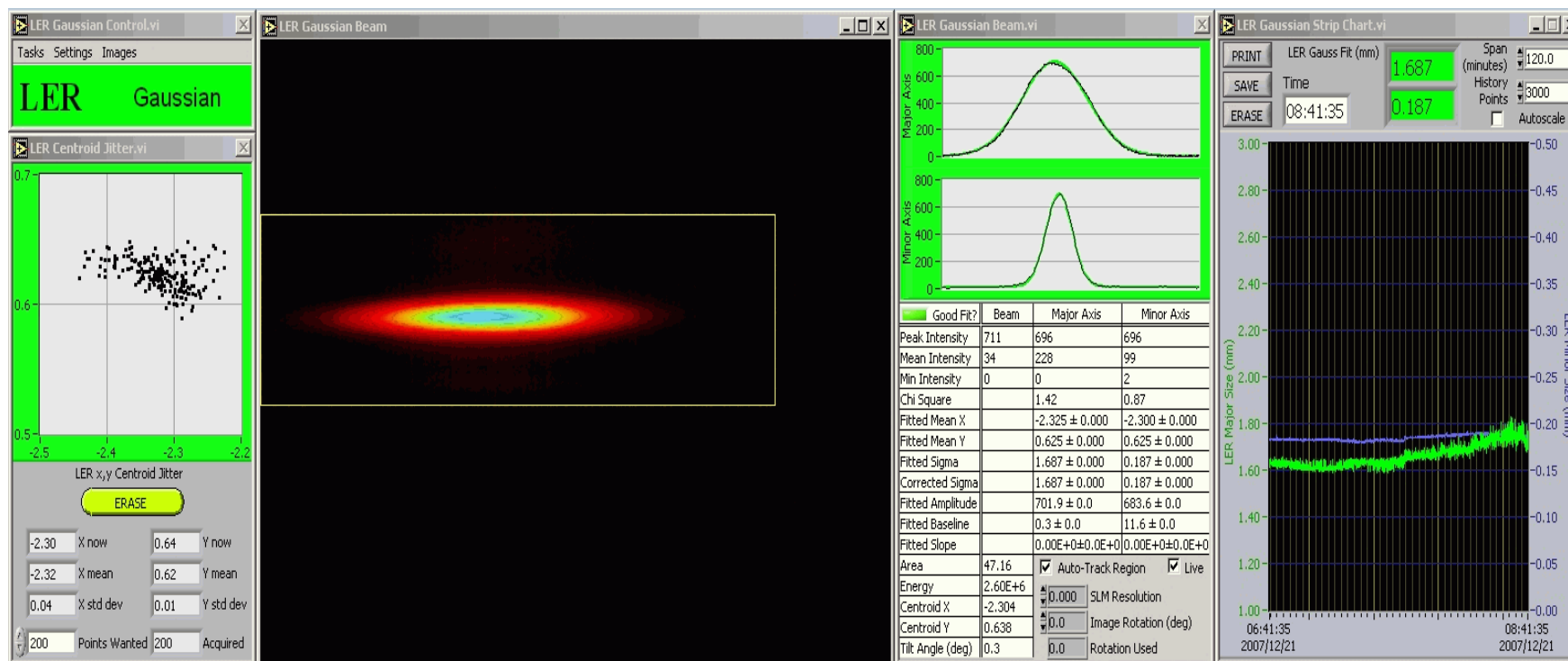
EPICS, MATLAB



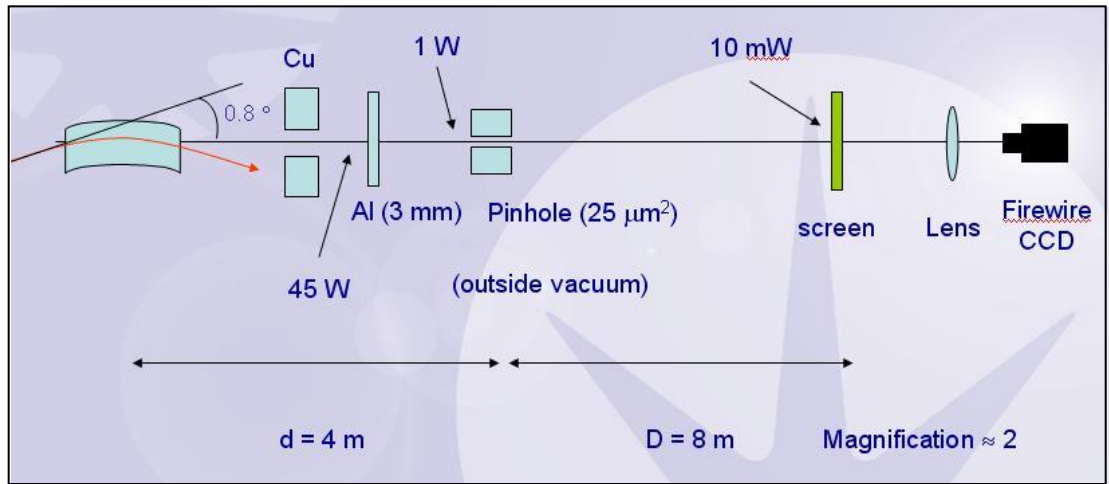
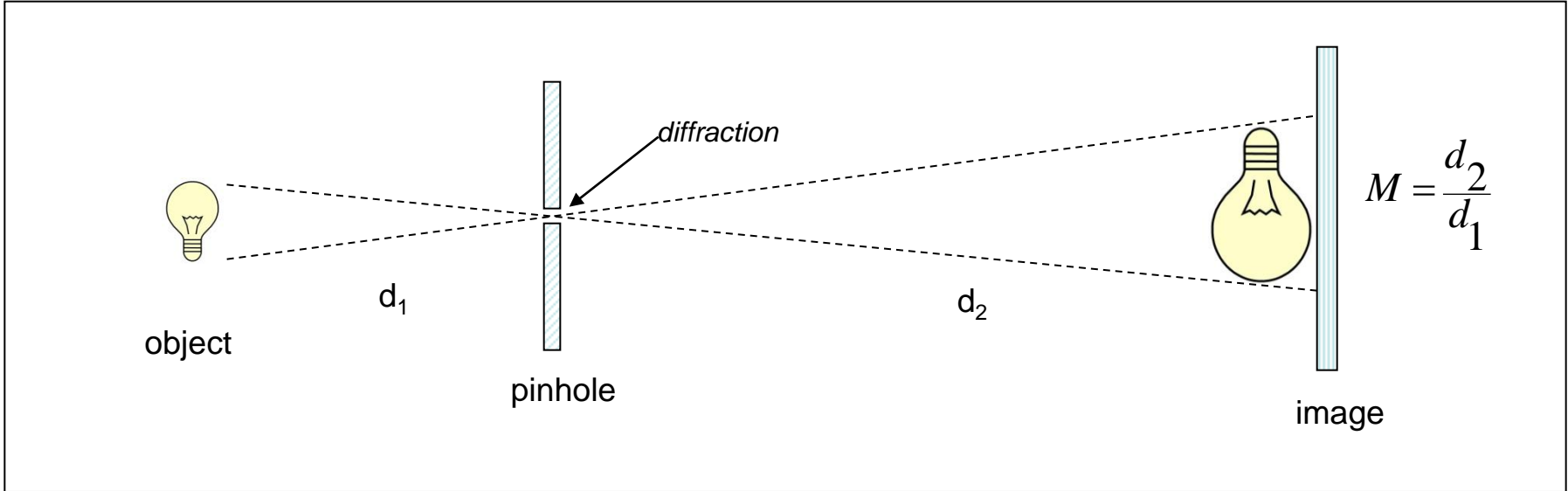
Beamspot at XSR



# PEP-II: visible light monitor software (LabView)

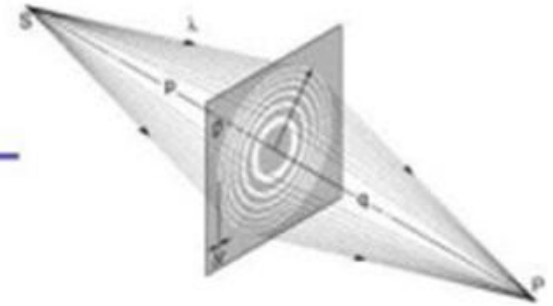


# X-ray pinhole cameras - Reduce diffraction with small $\lambda$





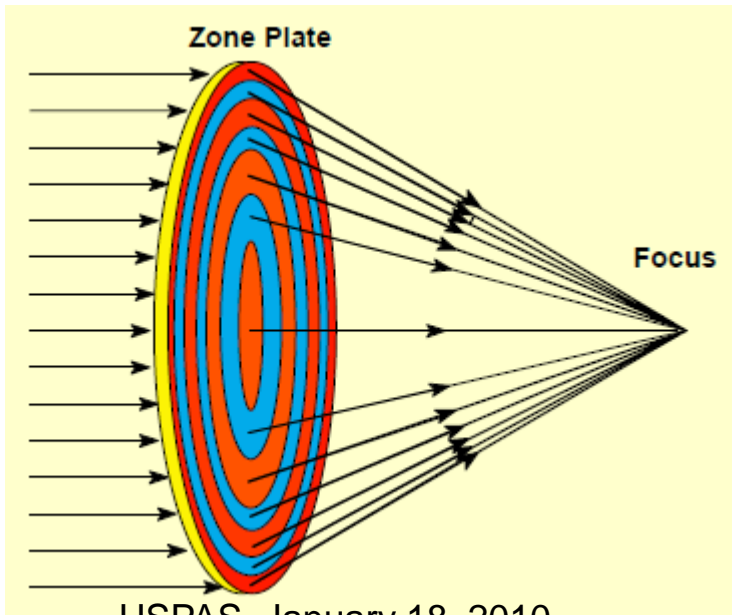
# Zone plate optics



➤ Zone plates act as a thin lens, with

$$f = 4N(\Delta r)^2 / \lambda$$

where  $N$  is the number of zones,  $\Delta r$  is the width of the outermost zone, and  $\lambda$  is the wavelength of the light. Thus the zone plate can act as a *linear monochromator* if one selects a particular focus (using a pinhole).



$n^{\text{th}}$  zone radius:

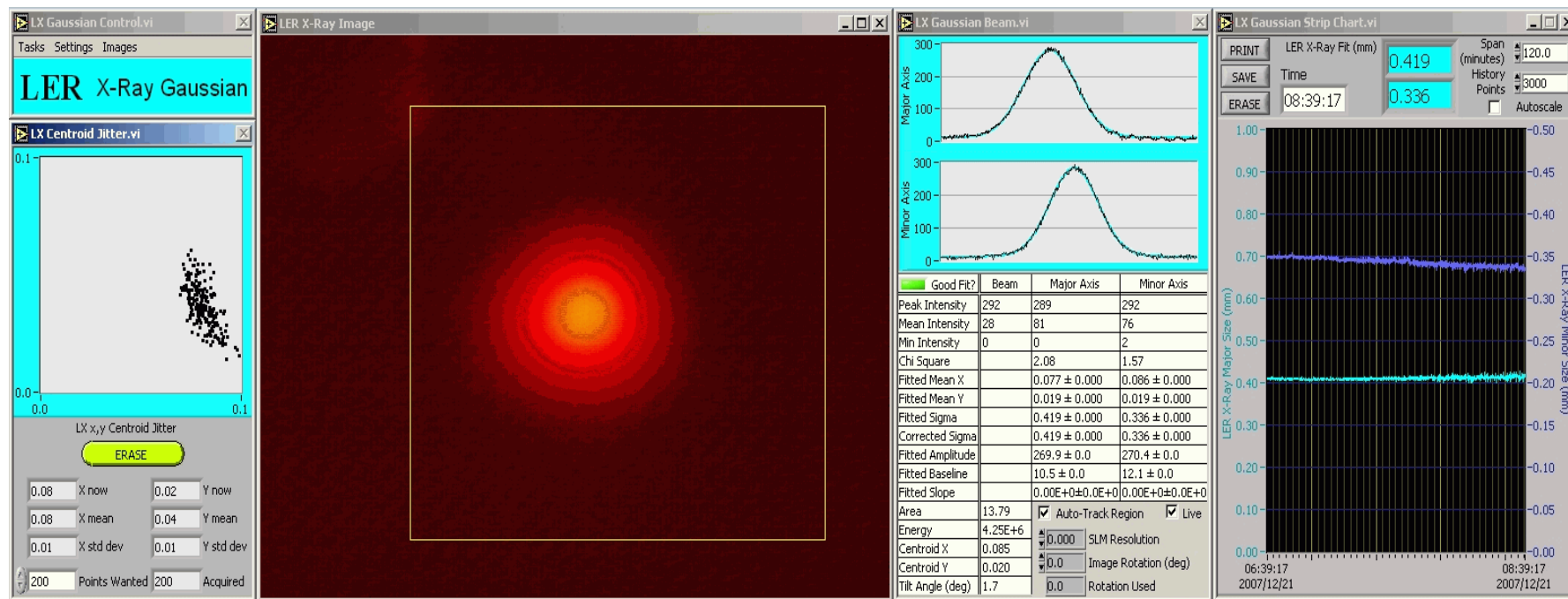
$$r_n = \sqrt{nf\lambda + \frac{n^2 \lambda^2}{4}} \approx \sqrt{nf\lambda}$$

$\lambda$  – photon wavelength

$f$  – focal length for the wavelength



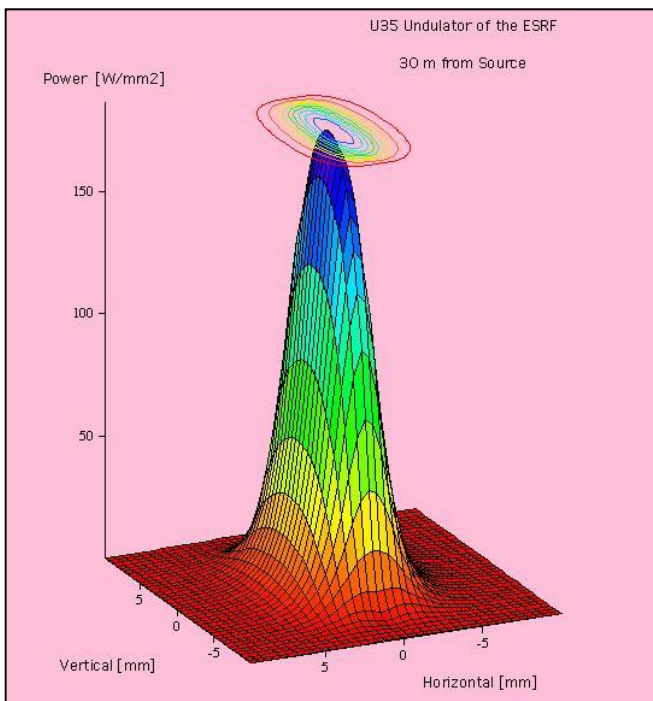
# PEP-II: x-ray pinhole camera software (LabView)





# Photon beam propagation programs

SRW  
(synchrotron radiation workshop)



Zemax  
(commercial product)

Surf. #	Type	Comment	Radius	Thickness	Glass	Semi-Diameter	Conic	Par 0 (unused)	Par 1 (unused)	P
007	Standard		Infinity	9630.000000		0.000000	0.000000			
ST0	Standard		Infinity	0.000000		17.000000	0.000000			
2	Coord Break			0.000000		-	0.000000			0.000000
3	COS_MXR DLL		Infinity	0.000000	MIRROR	109.896587	0.000000			0.000000
4	Coord Break			-3750.000000		-	0.000000			0.000000
5	Standard		Infinity	-1120.000000		23.931464	0.000000			
6	Coord Break			0.000000		-	0.000000			0.000000
7	Standard		Infinity	0.000000	MIRROR	26.162427	0.000000			
8	Coord Break			760.000000		-	0.000000			0.000000
9	Coord Break			0.000000		-	0.000000			0.000000
10	Coord Break			0.000000		-	0.000000			0.000000
11	Standard		Infinity	0.000000	MIRROR	27.546136	0.000000			
12	Coord Break			-3100.000000		-	0.000000			0.000000
13	Standard		Infinity	0.000000		33.117965	0.000000			
14	COS_MXR DLL	3860.829007 V		2000.000000	MIRROR	33.117702	0.000000			0.000000
15	Coord Break			0.000000		-	0.000000			0.000000
16	Standard	327.575864 V		-4000.000000	MIRROR	2.502201	0.000000			0.000000
IMA	Standard		Infinity			2.399420	0.000000			

Spot Diagram 1

3D Layout 1

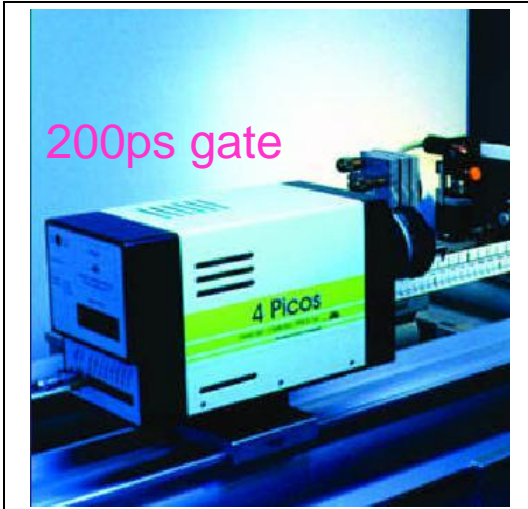
3D Layout 2

LENS #1: NO TITLE  
 OBJ #1: 2000 UNITS ARE MICRONS:  
 FOCUS: 4.000 1.000 1.120  
 OBJ #2: 0.200 9.000 10.370  
 SCALE: 1.000 REFERENCE: CHIEF RAY



# Cameras Part-II: Gated ICCD's

Stanford 4 Picos



Roper/PiMax

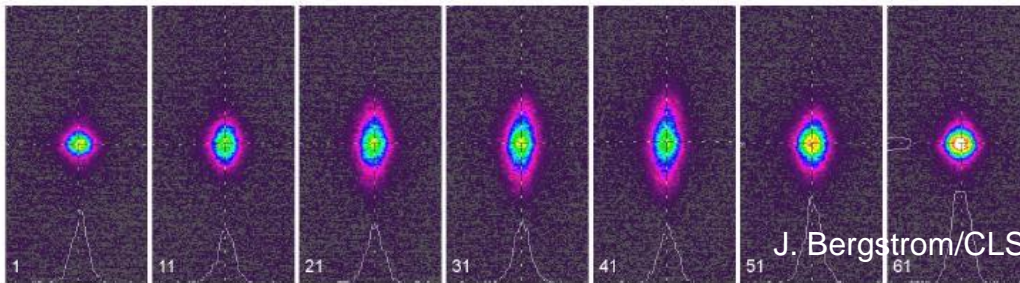
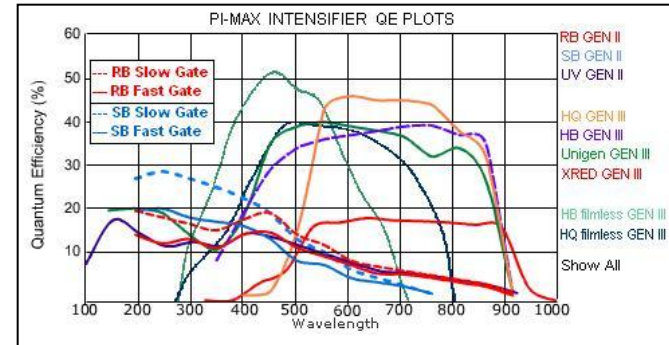
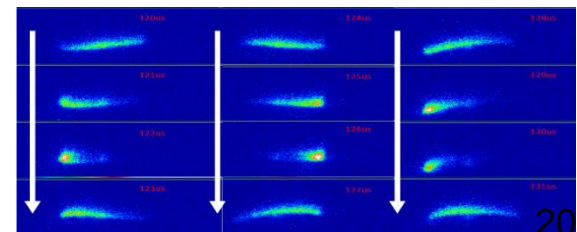


Fig. 8: Walking along a bunch train with the ICCD camera. This sequence shows every

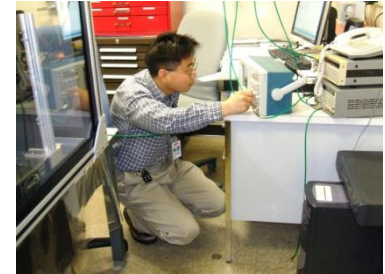
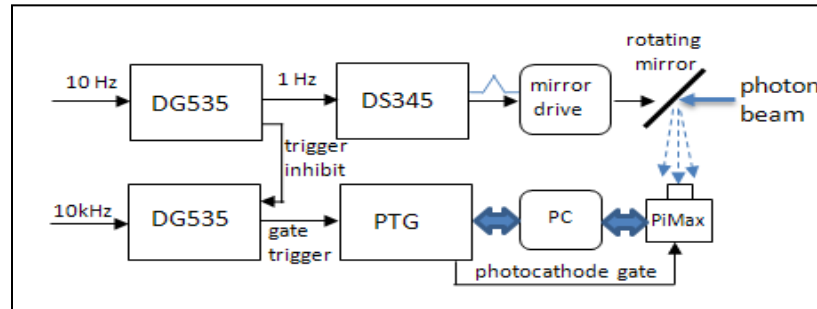


Injected beam in SPEAR3

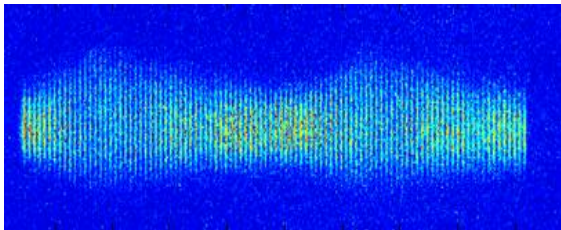




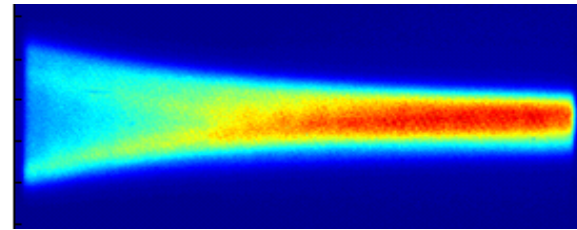
# The Roper/PiMax Camera Experiment



**Rotating Mirror and Timing System**



**Fast image sweep across photocathode**

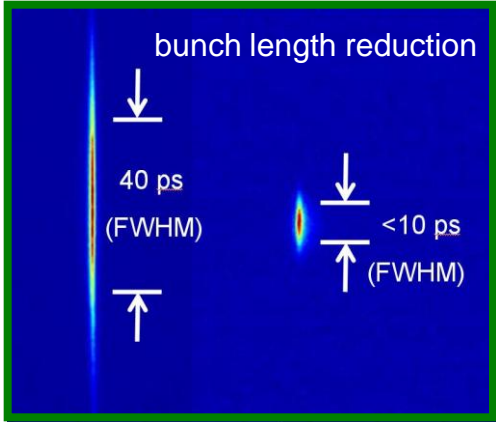


**Slow image sweep across photocathode**

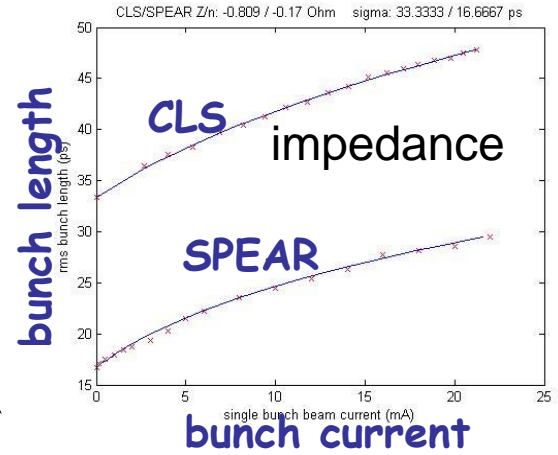


# Cameras III - Streak tubes

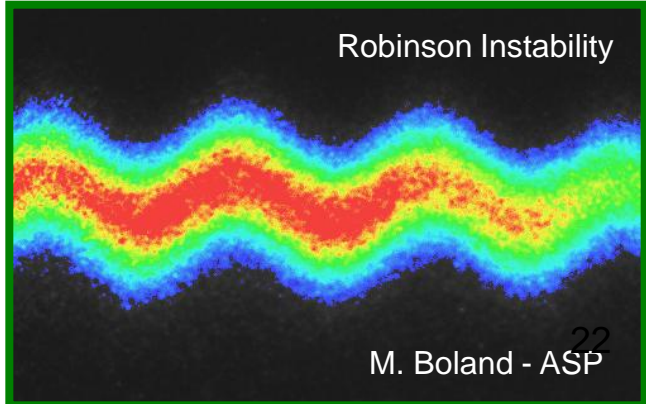
Hamamatsu C5680  
Optronis (ASP)



visible beam

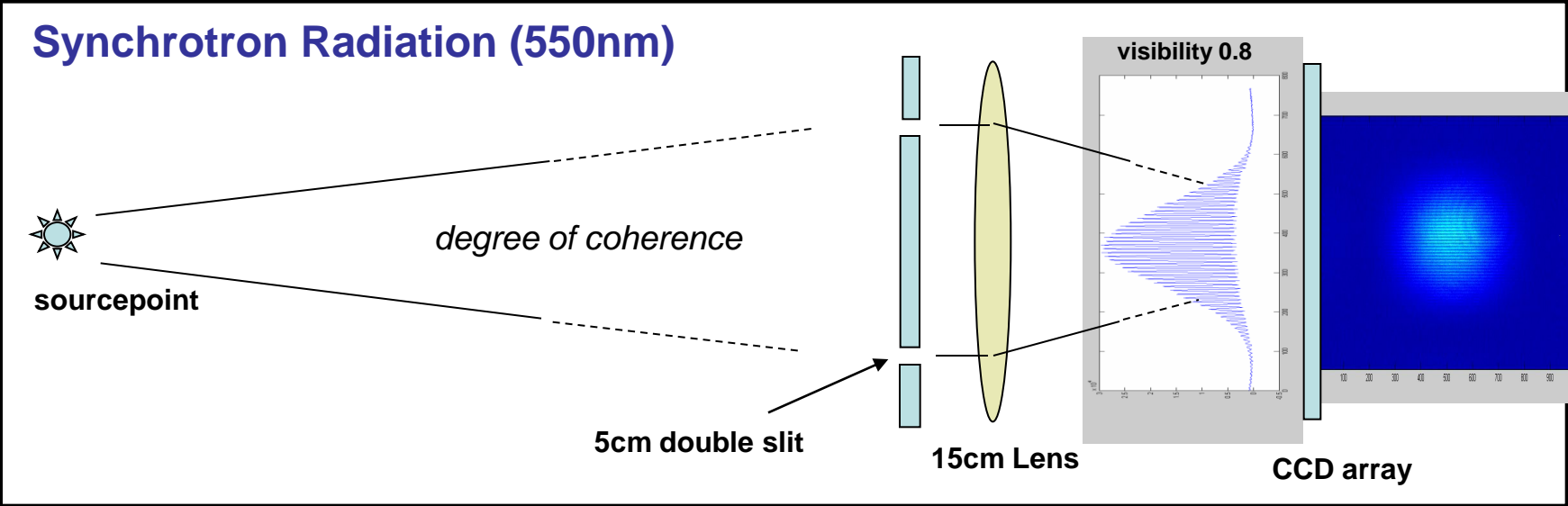
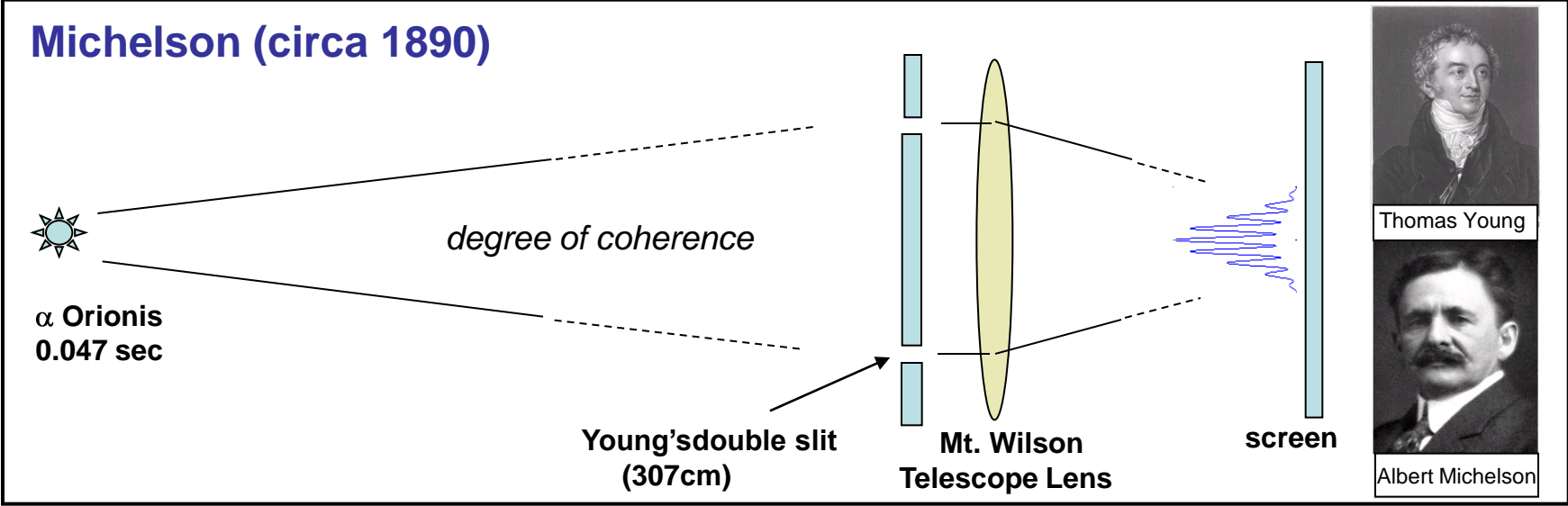


speed: up to 2 pixel/ps  
chromaticity: BP filter needed  
-bunch length  
-impedance and instabilities



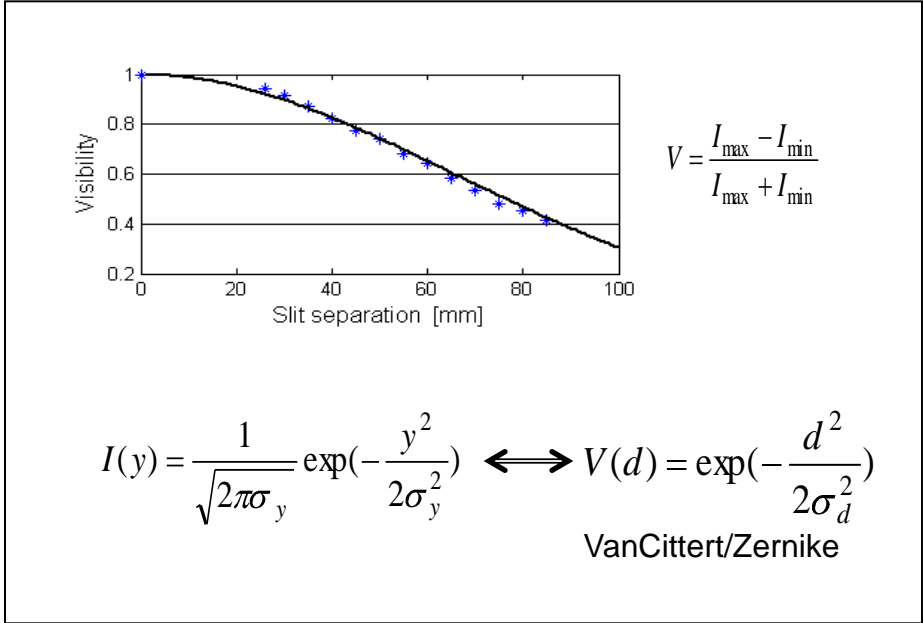
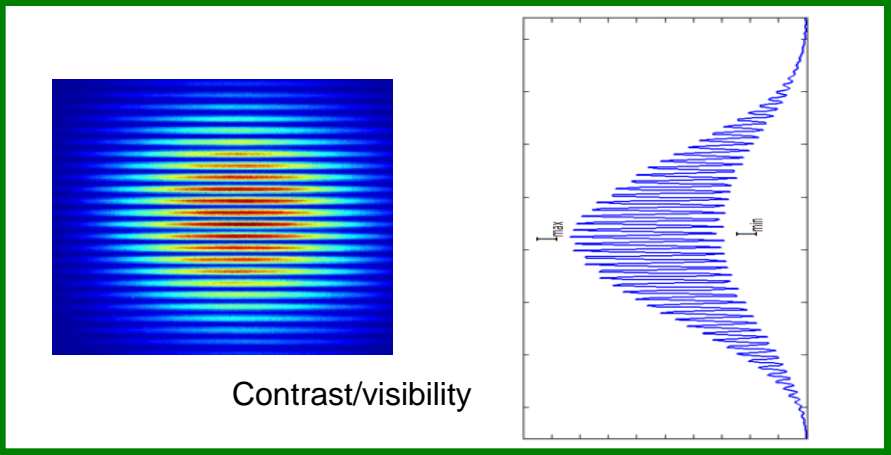
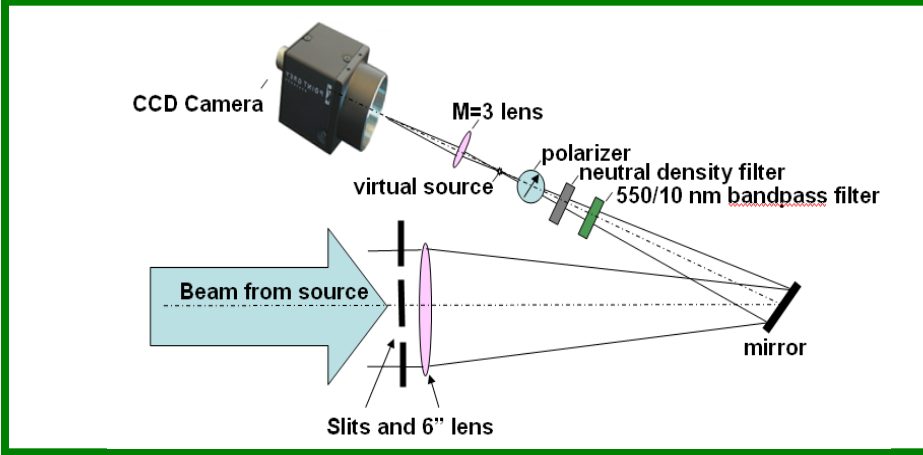


# The Stellar Interferometer





# The Stellar Interferometer (cont'd)



## Fringe Formulas

The slits have width  $a$  and center-to-center spacing  $d$ . The pattern from a single slit is:

$$I_{\pm}(\theta) = \left[ \frac{\sin\left[\frac{ka}{2}\left(\theta \mp \frac{d}{2s_0}\right)\right]}{\frac{ka}{2}\left(\theta \mp \frac{d}{2s_0}\right)} \right]^2$$

The interference from both slits at height  $y$  on the CCD, integrated over the optical bandpass filter, shows decreasing modulation with beam size:

$$I(y) = \int_{-\infty}^{\infty} \left[ I_+ + I_- + 2\sqrt{I_+ I_-} \exp\left(-\frac{(kd\sigma_y)^2}{2s_0^2}\right) \cos\left(\frac{kdy}{f + \Delta z}\right) \right] g(\lambda) d\lambda$$

The ray leaving the slit at an angle  $\theta$  hits the CCD at height  $y$ :

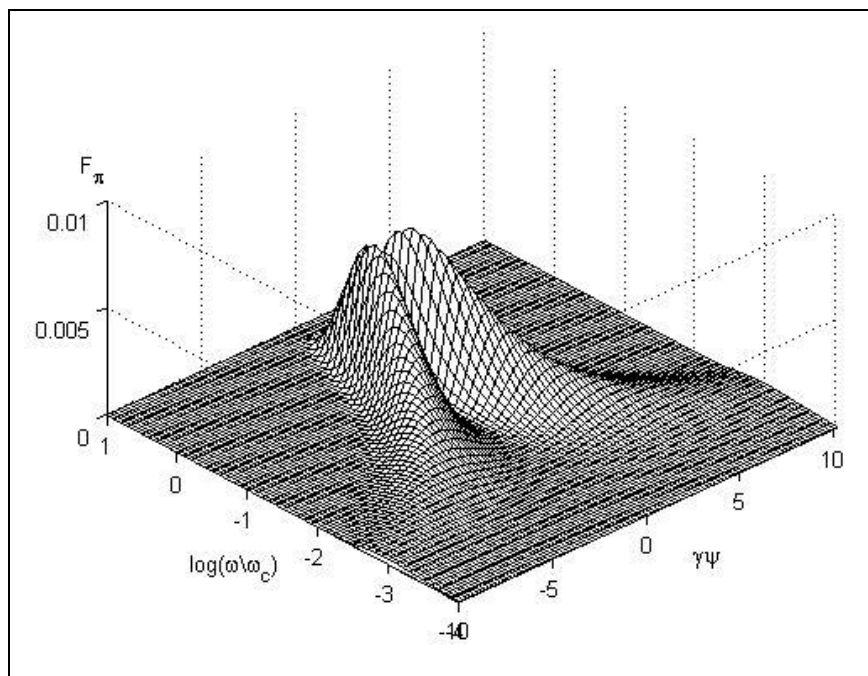
$$\theta_{\pm}(y) = \frac{y \pm \frac{d\Delta z}{2f}}{f + \Delta z \left(1 - \frac{\Delta s}{f}\right)}$$

A. Fisher - SLAC

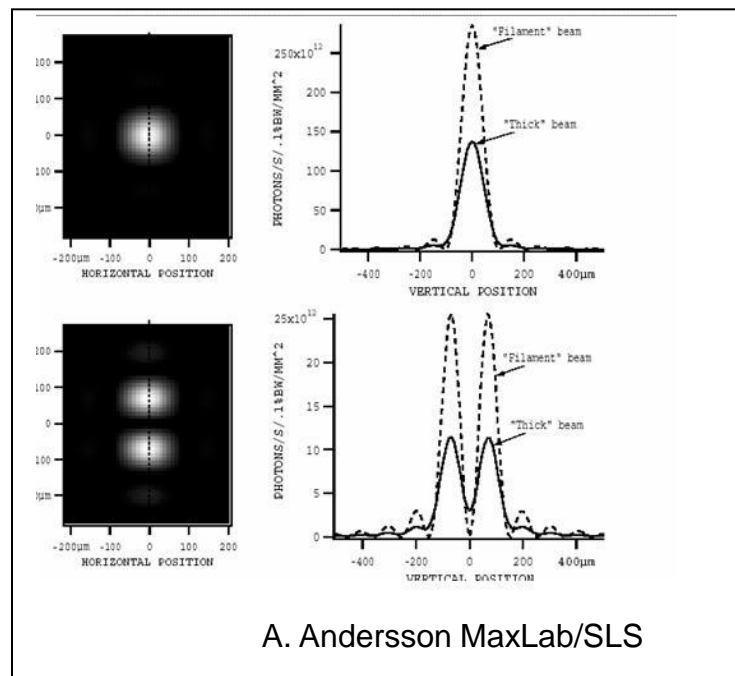


# Small beam size - Vertical polarization technique

Vertical angular spectral density

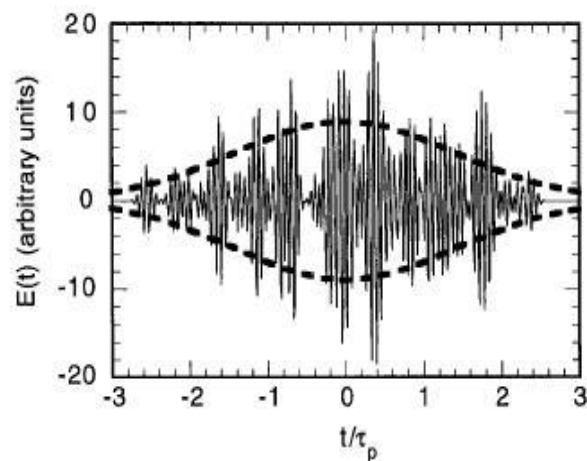


Measurements at MaxLab



# Bunch Length Measurement - Statistical Fluctuations

Intra-pulse fluctuation of the electric field  
 Statistics of pulse-to-pulse variations



$$\delta^2 = \frac{\sigma_W^2}{\langle W \rangle^2} = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} dt dt' \frac{|K(t-t')|^2}{|K(0)|^2} I(t) I(t')$$

$$\delta^2 = 1/\sqrt{1 + 4\sigma_t^2 \sigma_\omega^2}$$

$$\frac{\text{Signal}}{\text{Noise}} \sim \frac{1}{\sqrt{M}}$$

$$\frac{\langle N \rangle}{\sigma} \sim \left( \frac{\tau_{\text{coh}}}{\tau_{\text{pulse}}} \right)^{1/2}$$

Fig. 1. Electric field of a pulse of incoherent radiation as a function of time. The ratio  $\Delta\omega/\omega_0 = 0.1$ , and the parameter  $N = 10$ . The dashed lines show  $\sqrt{I(t)}$ .

Hofmann: chapter 15

G. Stupajkov/SLAC

# Fluctuation measurement

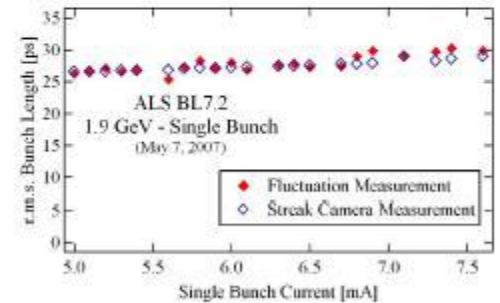
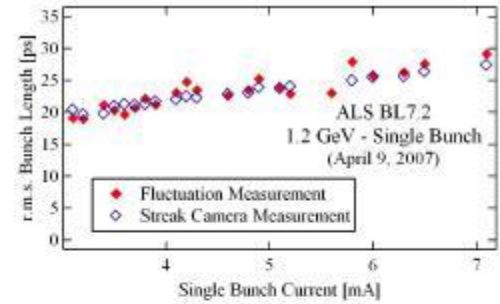
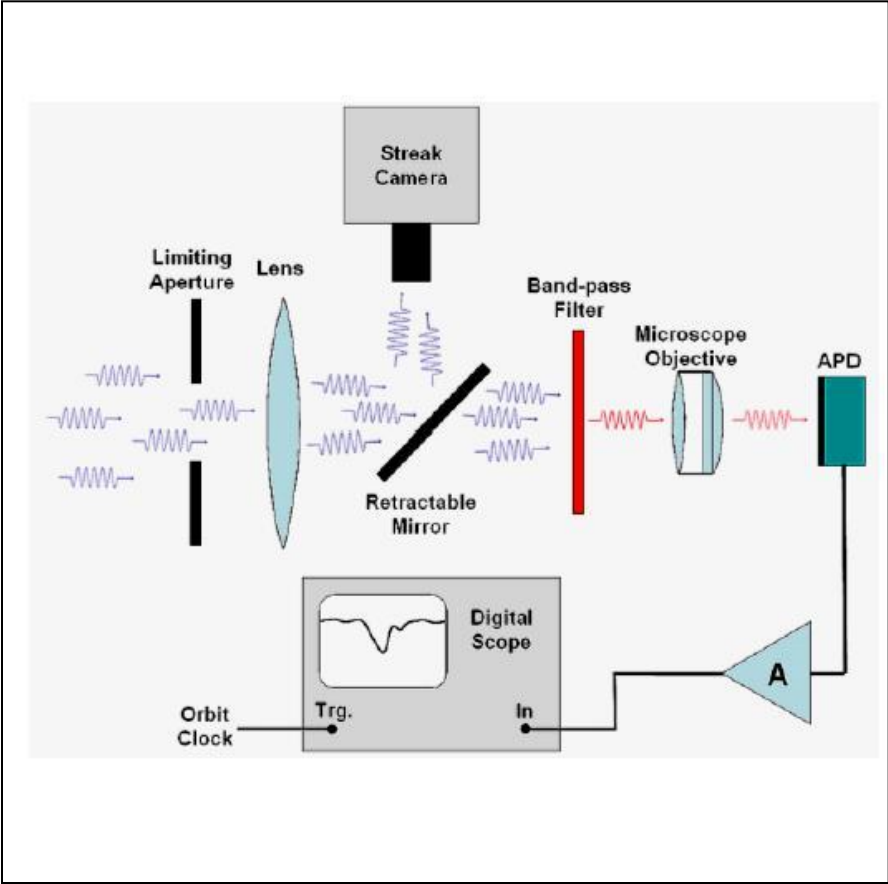


Figure 3: Examples of fluctuation and streak-camera bunch length measurements at the ALS for different beam parameters.

F. Sannibale/LBL



## Summary of beam size measurements

- Photon emission provides valuable diagnostic of  $e^-$  beam
- Need to unfold  $\gamma_r$ , DOF, diffraction, PSF, etc. from image
- Visible has advantage of commercial optics and cameras but suffers from large  $\gamma_r$  and diffraction
- Broad array of cameras, fast shutters, streak frames
- X-ray pinhole has advantage of less diffraction but generally less versatile
- Interferometers and central-null technique improve resolution
- Fluctuation measurements cheaper than streak, provide insight
- other techniques:
  - screens, OTR, wires and lasers in transmission lines
  - scraper in storage ring (quantum lifetime)
  - response matrix analysis